

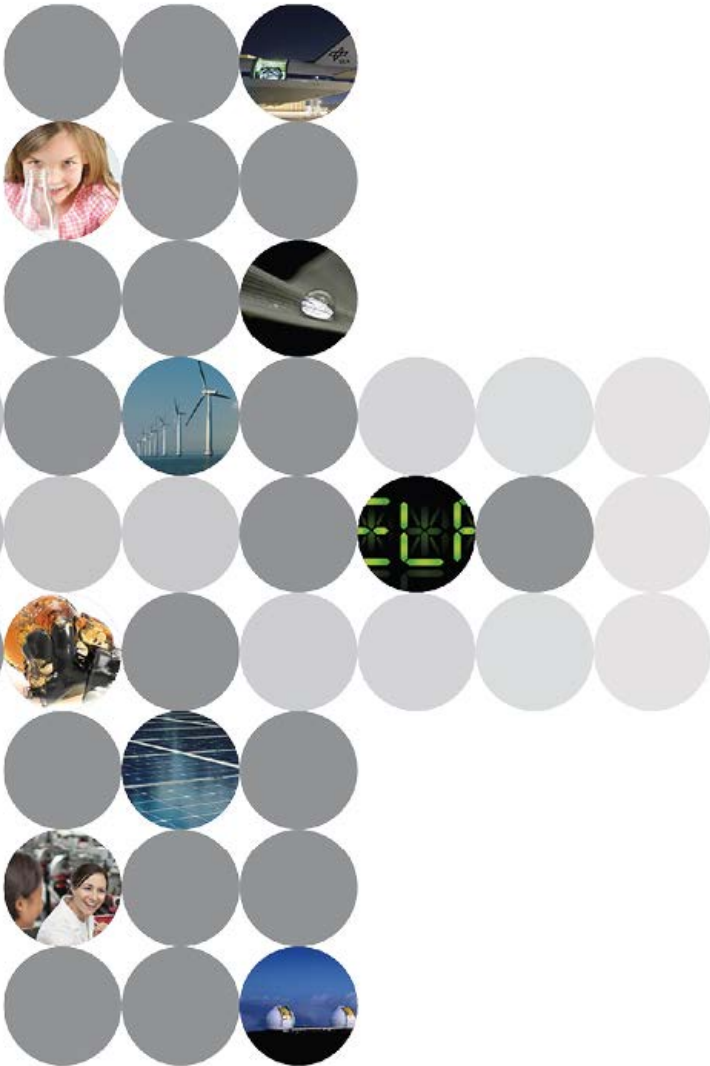
CSEM

*technologies
that make **the** difference*

Technologies lasers bas bruits et sources de bruit

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Bordeaux
18.01.2018*

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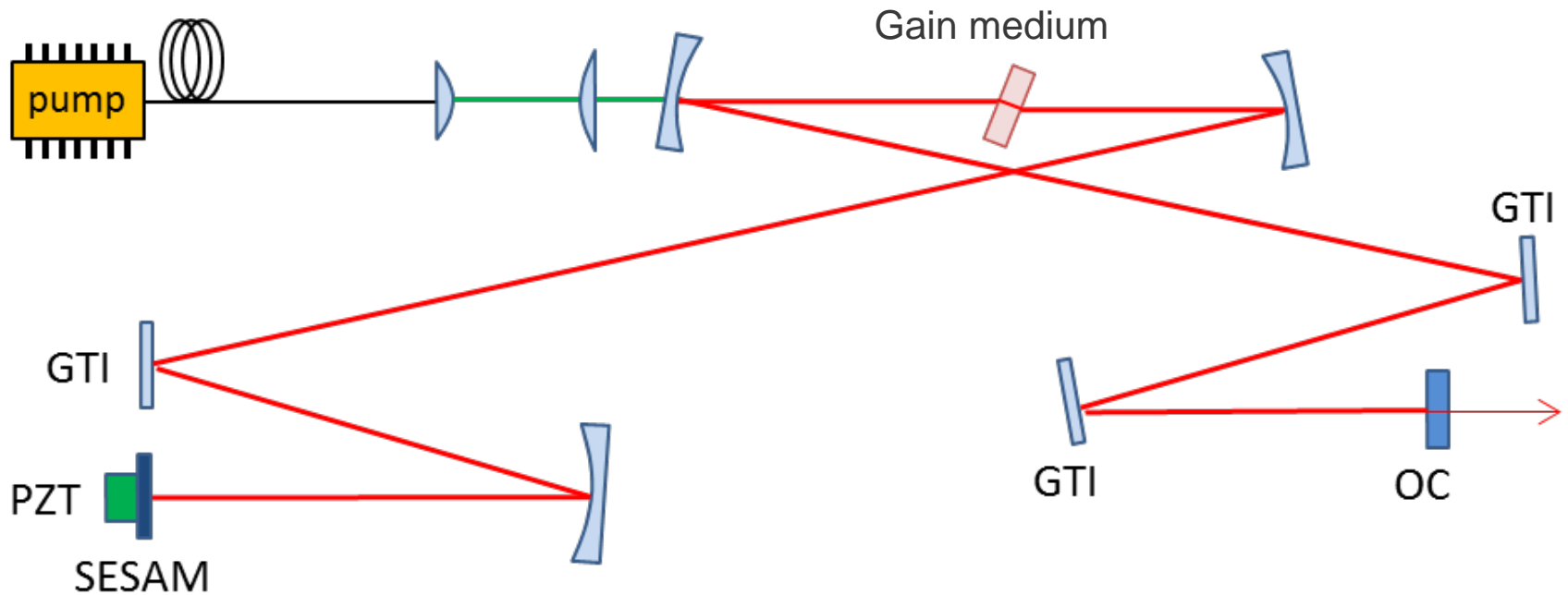


Outline

- :: Mode-locked lasers as low-noise oscillators
- :: Amplitude, CEO phase, optical phase, and timing jitter noise
- :: Elastic tape model & frequency dependance on gain and loss perturbations
- :: Noise sources (vibration, thermal & acoustics, pump RIN, quantum noise, dispersion, slow saturable absorber, self-steepening and Kramers-Kronig)
- :: Results (RIN, pulse jitter and microwave generation)
- :: Stabilization techniques for fceo and frep
- :: Summary and conclusions

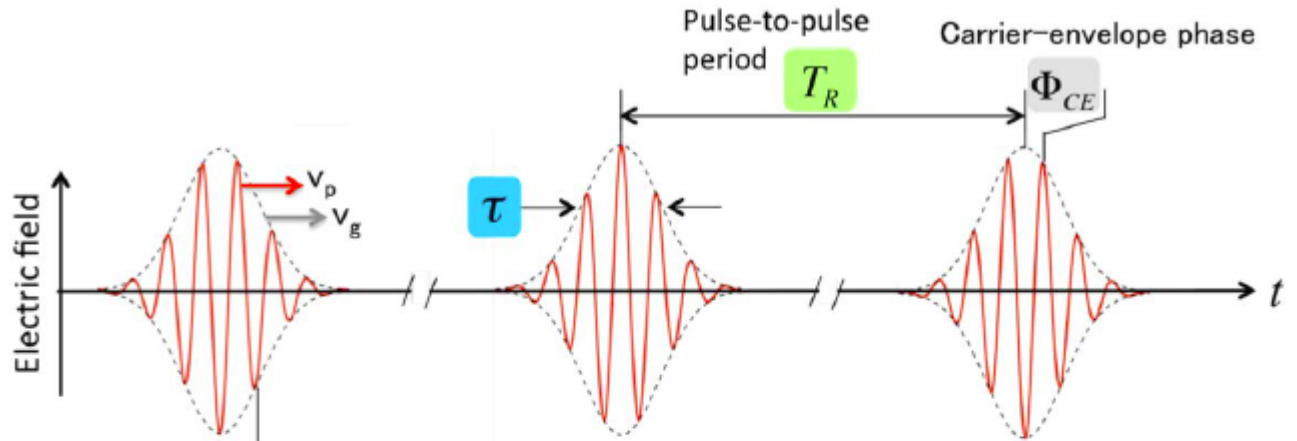
Mode-locked laser based on DPSSL technology

:: Diode-pumped solid-state laser - DPSSL



Output of a modelocked femtosecond laser

Time domain



Frequency domain

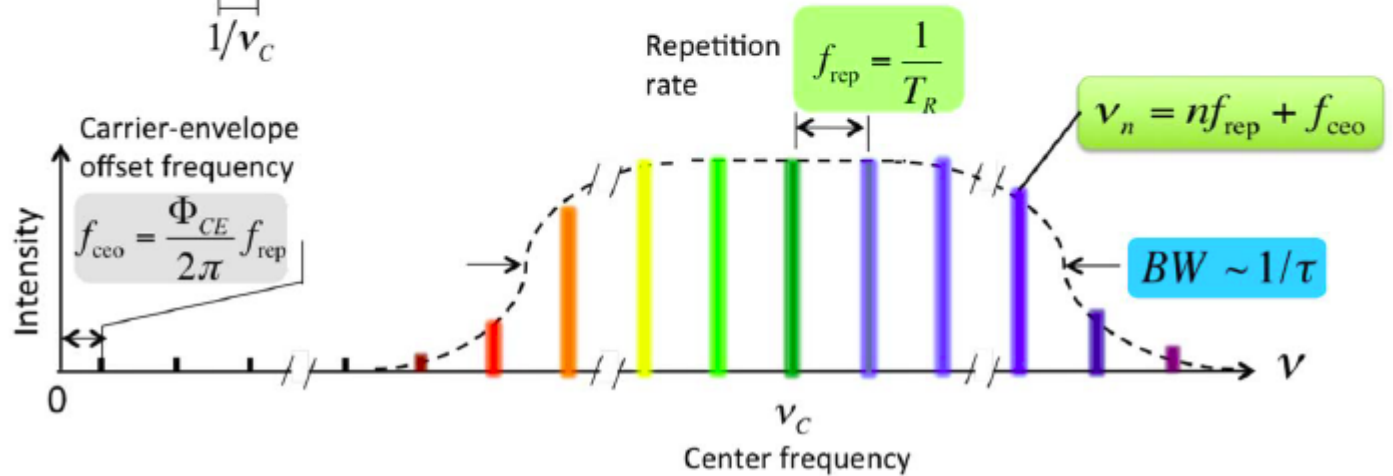
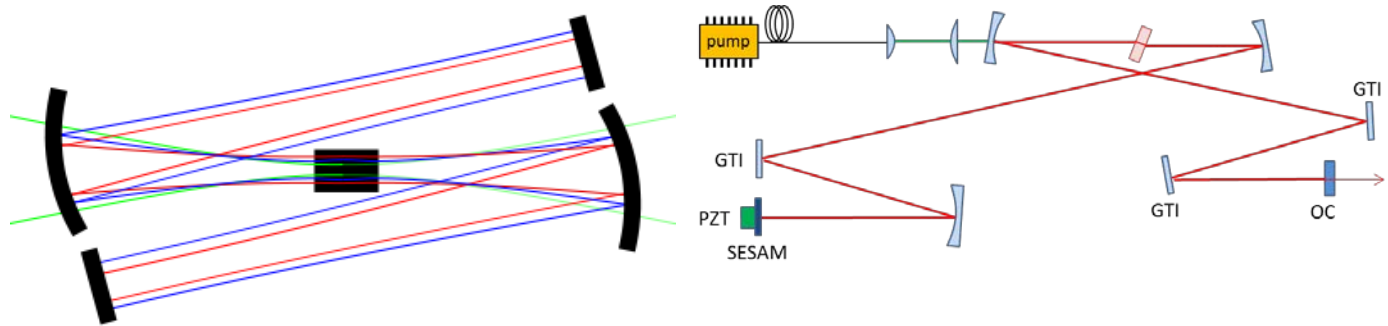


Figure from J. Kim and Y. Song, Adv. In Opt. And Phot. 8 465-540, 2016

Passive modelocking of solid-state lasers

:: Kerr lens and saturable absorber

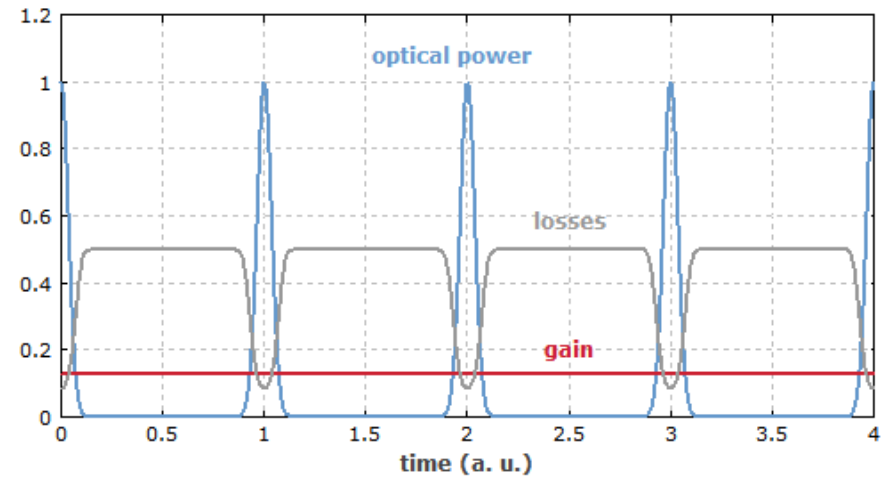


Parameter	Kerr lens	Saturable absorber
Recovery time	Fast (instantaneous)	«Slow»
Self-starting ML	Usually not	Yes
Wrt noise	Best case	The faster the better
Laser type	Ti:sapphire typically	Any type

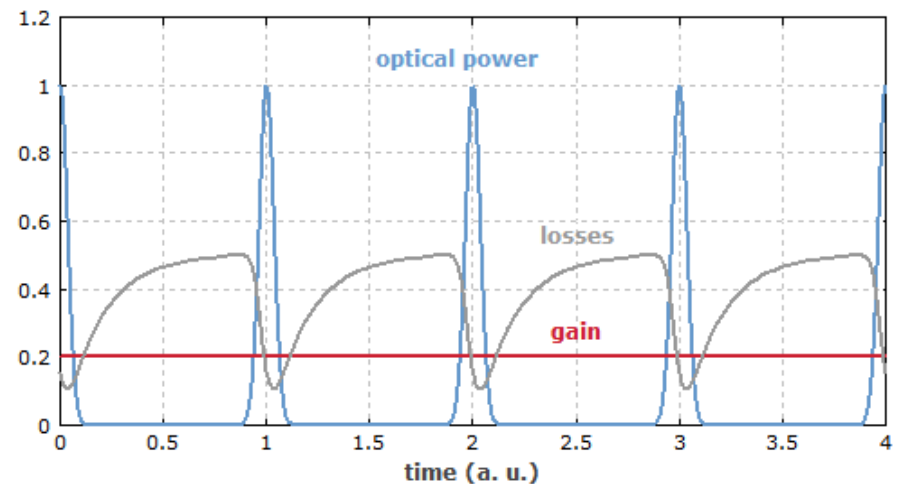
Left figure from wikipedia

Dynamics in solid-state laser vs. absorber speed

:: Fast saturable absorber

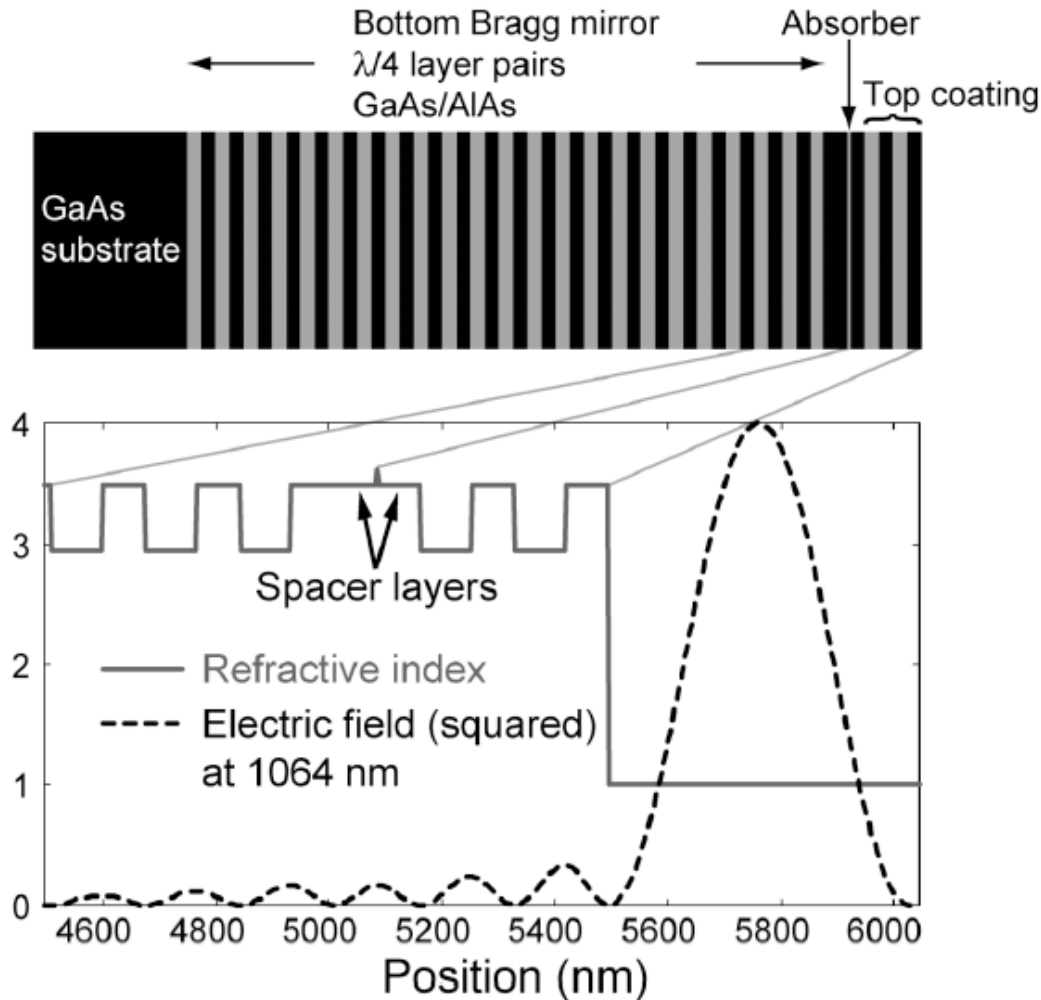


:: Slow saturable absorber



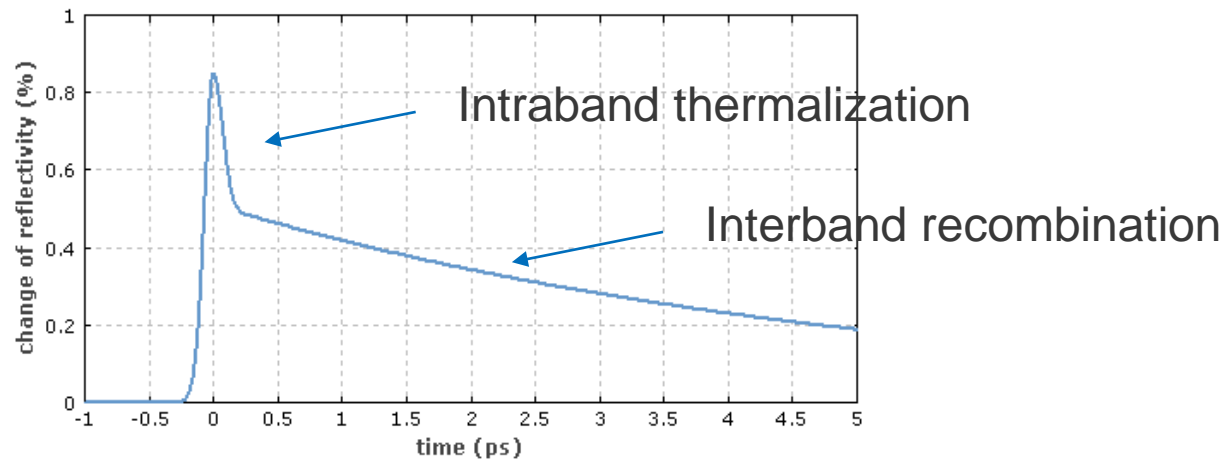
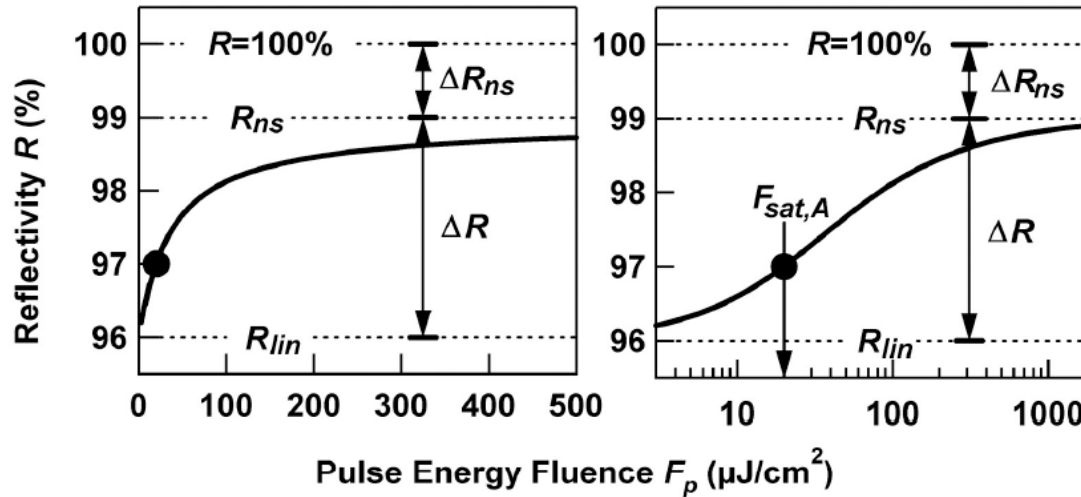
Figures from www.rp-photonics.com

Passive mode-locking with Semiconductor Saturable Absorber Mirror (SESAM)



Figures from S. Lecomte, PhD thesis

Passive mode-locking with Semiconductor Saturable Absorber Mirror (SESAM)



Figures from S. Lecomte, PhD thesis and www.rp-photonics.com

Relevant laser parameters wrt noise

- :: Amplitude – RIN (Relative Intensity Noise)
- :: Pulse train timing jitter
- :: Optical phase of the comb modes
- :: Carrier-envelope offset phase noise

$$A(t) = [A_0 + \Delta A_0(t)] \sum_{m=-\infty}^{+\infty} a(t - mT_R + \Delta T_R(t)) \exp[j\{2\pi\nu_c t + m\Phi_{CE} + \Delta\theta(t)\}]$$

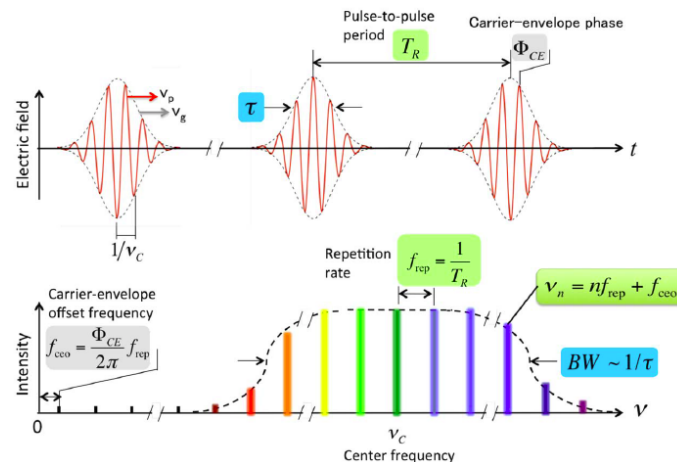
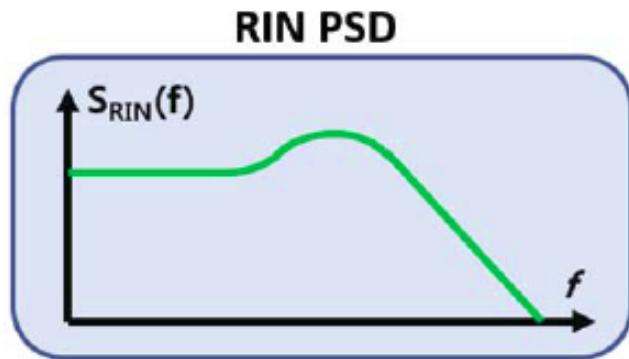
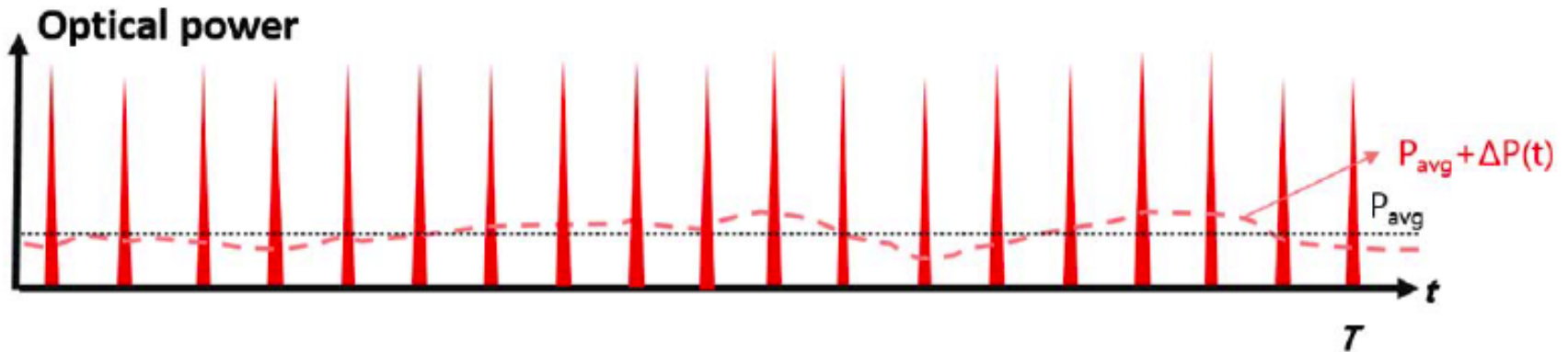


Figure from J. Kim and Y. Song, Adv. In Opt. And Phot. 8 465-540, 2016

Amplitude noise - RIN



$$RIN = \frac{\langle \Delta P(t)^2 \rangle_T}{\langle P(t) \rangle_T^2}$$

$$\lim_{T \rightarrow \infty} \frac{1}{T} \left\langle \left| \int_{-T/2}^{+T/2} \frac{\Delta P(t)}{P_{avg}} e^{-j2\pi ft} dt \right|^2 \right\rangle$$

$$\int_{-f_{rep}/2}^{f_{rep}/2} S_{RIN}(f) df$$

Quantum limit:

$$S_{RIN}^{shot\ noise}(f) = \frac{2h\nu_c}{P_{avg}}$$

Pulse train timing jitter

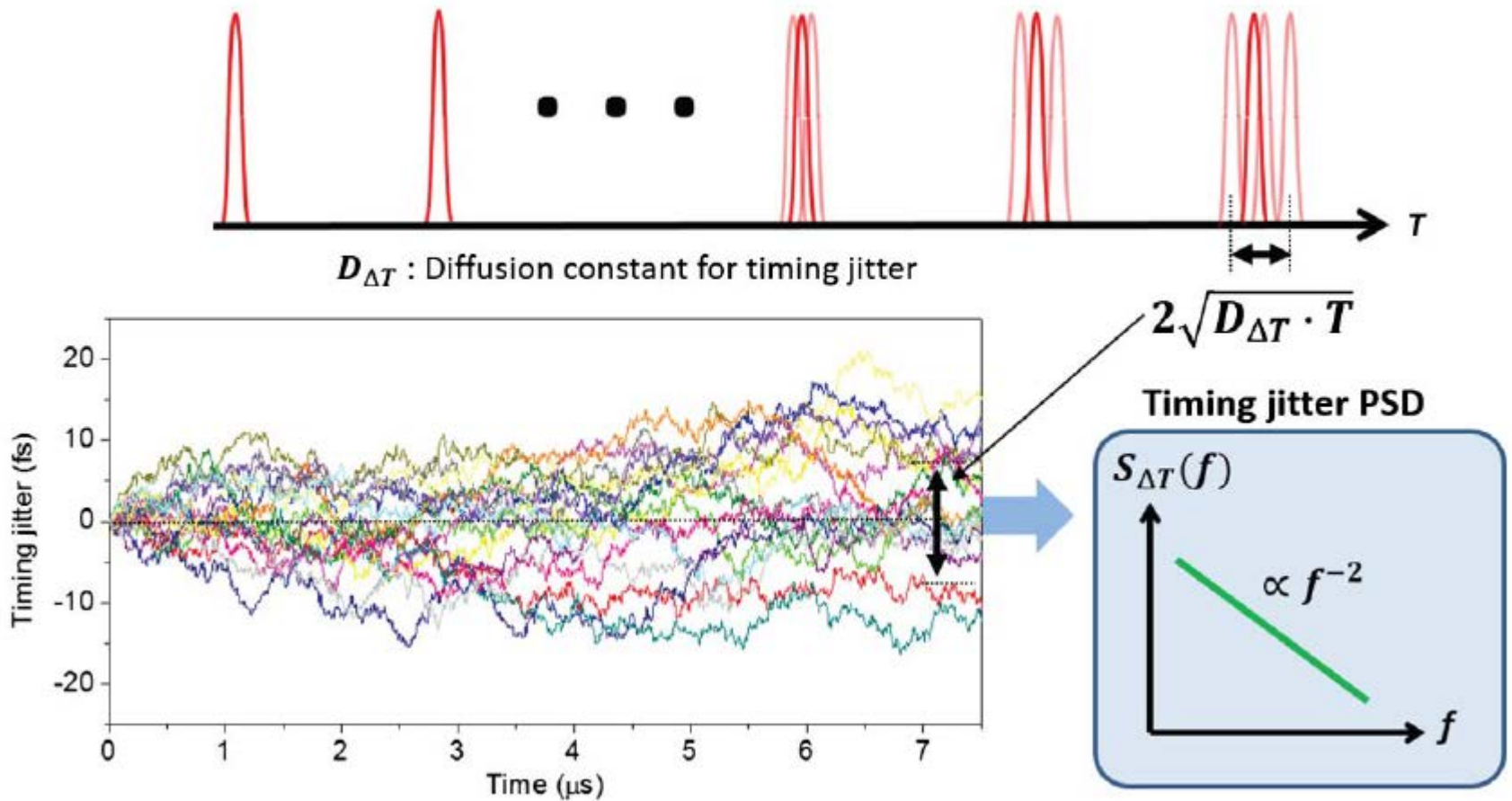


Figure from J. Kim and Y. Song, Adv. In Opt. And Phot. 8 465-540, 2016

Optical phase (comb-mode linewidth)

In case of ASE quantum limit and uncorrelated noise:

$$\Delta\nu_n = \Delta\nu_{\Delta\theta} + [2\pi\tau(\nu_n - \nu_c)]^2 \Delta\nu_{\Delta T}$$

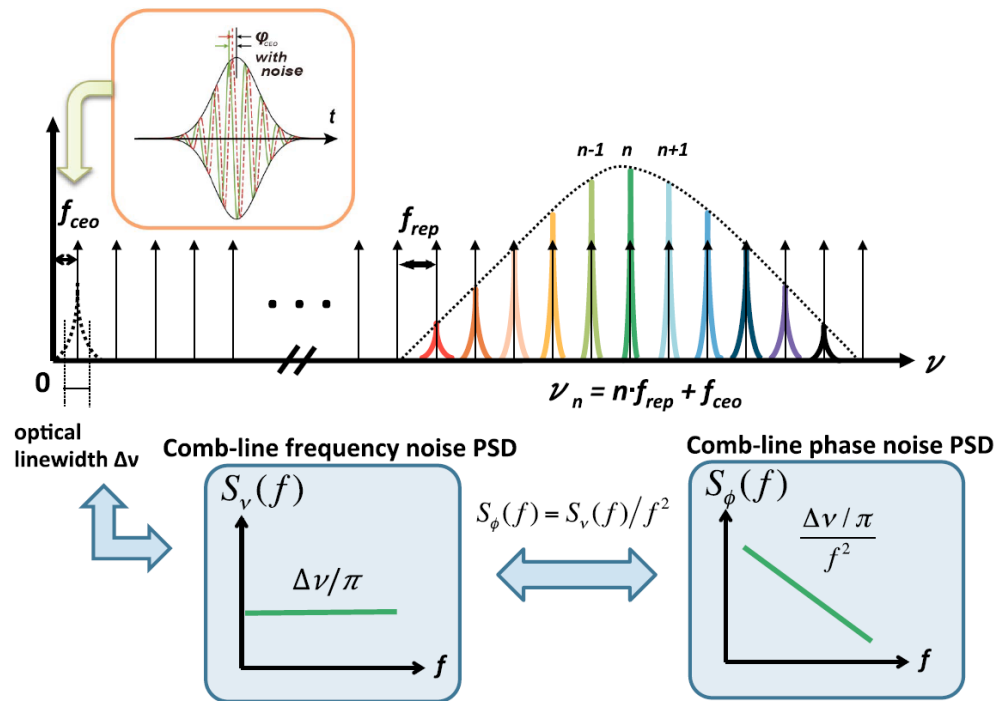
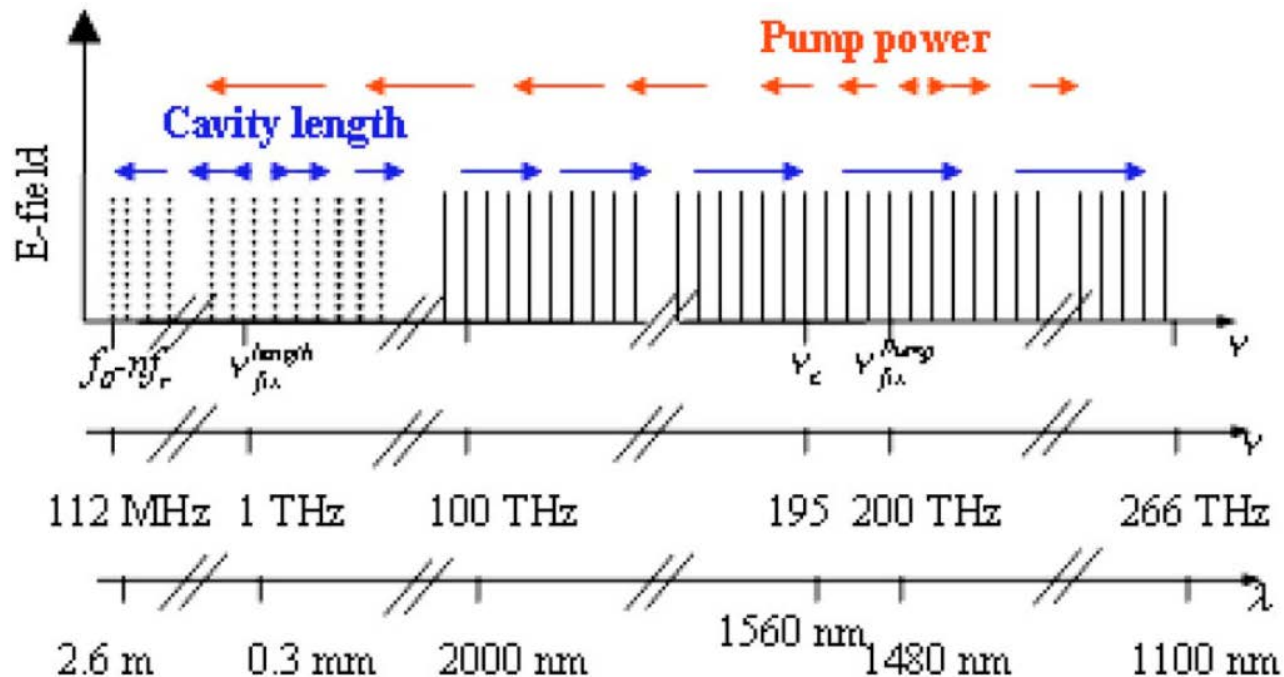


Figure from J. Kim and Y. Song, Adv. In Opt. And Phot. 8 465-540, 2016

Elastic tape model

- ∴ Influence of noise sources depends on fix frequency (of the perturbation) and on considered comb mode of interest



From N. R. Newbury and W. C. Swann, JOSA B 24 1756-1771, 2007

Elastic tape model – frequency dependance

:: Frequency dependance is defined by the physics of the laser

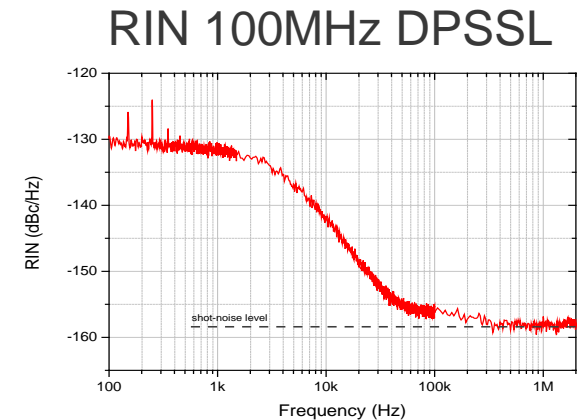
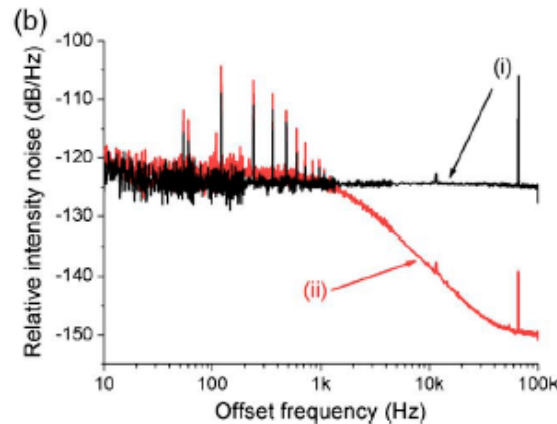
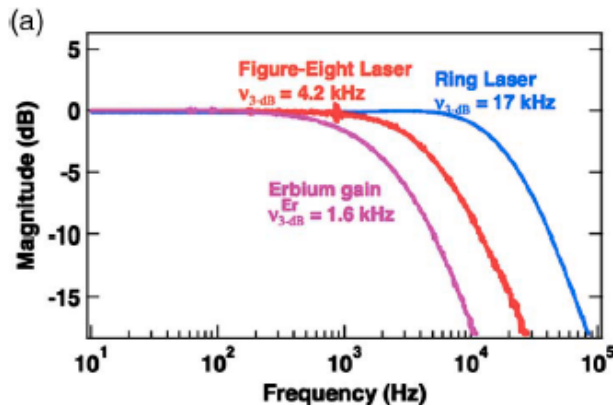
- Gain modulation bandwidth:

$$\omega_1 \approx T_G^{-1} \left(\frac{\partial q}{\partial E_p} E_p \right)^{-1}$$

- Loss modulation bandwidth:

$$\omega_2 \approx T_R^{-1} \left(\frac{\partial q}{\partial E_p} E_p \right)^{-1}$$

$$\omega_2 \gg \omega_1$$



From C.-C. Lee et al. Opt. Lett. 37, 3084-3086, 2012 and from J. Kim and Y. Song, Adv. In Opt. And Phot. 8 465-540, 2016

Elastic tape model

Table 1. Fixed Point, Frequency Dependence, and Magnitude of the Various Contributions to the Frequency Noise on the Comb Lines

Noise Term	Fixed Point (ν_{fix})	Frequency Dependence	Magnitude at $f=1$ Hz $S_r(1)$ in Units of 1/Hz	Suppress by
Environmental (length) ^{a,b}	0–3 THz (1 THz)	f^{-1}	10^{-22}	Environmental isolation
Environmental (loss or pump power)	ν_c (200 THz)	$\{f(1+(f/f_{3\text{ dB}})^2)\}^{-1}$ $f_{3\text{ dB}}=5-10\text{ kHz}^c$	10^{-21}	Environmental Isolation, reduce $f_{3\text{ dB}}$
Pump noise	$\sim \nu_c$ (200 THz)	$\{1+(f/f_{3\text{ dB}})^2\}^{-1}$ $f_{3\text{ dB}}=5-10\text{ kHz}$	3×10^{-24}	Reduce pump RIN and cavity $f_{3\text{ dB}}$
Intracavity ASE (quantum limit)	$\sim \nu_c$ (190 THz)	f^0	3×10^{-25}	Reduce effective cavity loss
Supercontinuum and shot noise ^d	NA ^c	f^2	6×10^{-23} to 6×10^{-24}	Higher peak powers
Environmental (external path length) ^e	0–2 THz	f	10^{-32}	Minimize extra path lengths

From N. R. Newbury and W. C. Swann, JOSA B 24 1756-1771, 2007

Comb-line frequency noise

Effect	Frequency noise Power Spectral Density (Hz ² /Hz)	Physical cause
Quantum noise (QL)	$S_{\nu}^{\text{ST}} = \frac{\Theta h \nu_c l_{\text{tot}}}{8\pi^2 P_{\text{int}}} \cdot f_{\text{rep}}^2$	Contribution of losses and gain to comb-line
ASE induced timing drift	$S_{\nu,n}^{\text{ASE}}(f) = S_{\nu}^{\text{ST}} + (\nu_n - \nu_c)^2 \cdot S_{f_{\text{rep}}}^{\text{ASE}}(f)/f_{\text{rep}}^2$	For any comb line n
Pump RIN induced*	$S_{\nu,n}^{\text{Pump}} = (n - n_{\text{fix}}^{\text{Pump}})^2 S_{f_{\text{rep}}}^{\text{Pump}}$	
Length jitter induced	$\nu_{\text{fix}}^{\text{length}} = \nu_c (1 - v_g^L/v_p^L)$	

$$* : \frac{df_{\text{ceo}}}{dP} = \frac{\beta_0}{2\pi} \left(\frac{df_{\text{rep}}}{dP} \right) + \frac{f_{\text{rep}}}{2\pi} \left(\frac{d\phi_0}{dP} \right) \quad \& \quad n_{\text{fix}}^{\text{Pump}} \equiv -(df_{\text{ceo}}/dP)/(df_{\text{rep}}/dP)$$

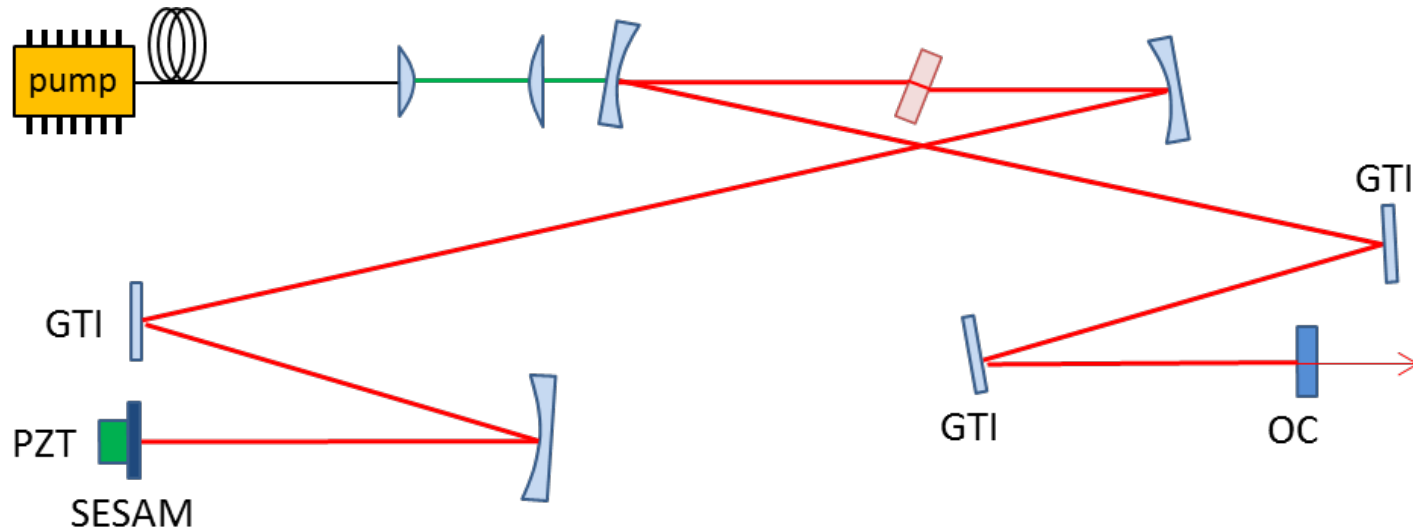
$$S_{f_{\text{rep}}}^{\text{ASE}}(f) = (2\pi f_{\text{rep}})^2 f^2 S_{\Delta T}^{\text{ASE}}(f)$$

Timing jitter

Effect	Jitter Power Spectral Density (fs ² /Hz)	Physical cause
Quantum noise (QL)	$S_{\Delta t}^{QL}(f) = 0.5294 \frac{1}{(2\pi f)^2} \frac{h\nu_0}{E_p} \frac{g}{T_{rt}} \tau_p^2$	Contribution of losses and gain to timing jitter. Fundamental quantum limit, no dispersion is considered.
Gordon-Haus (G-H)	$S_{\Delta t}^{G-H}(f) = \left(\frac{D_2}{f T_{rt}} \right)^2 \frac{g}{T_{rt} \tau_p^2} \frac{0.21}{(2\pi f)^2 + \left(0.21 \frac{g}{T_{rt}} \frac{1}{\tau_p^2 \Delta f_g^2} \right)^2}$	Pulse center optical frequency fluctuation due to quantum noise coupled to intracavity dispersion
Kramers-Kronig (K-K)	$S_{\Delta t}^{K-K}(f) = \left(\frac{1}{2\pi \Delta f_g} \right)^2 \left[1 + \left(\frac{E_p \Delta R}{2\pi f T_{rt}} \right)^2 \right] S_I(f)$	Coupling of variation of refractive index through inversion level and pump power fluctuations
Self-steepening (S-S)	$S_{\Delta t}^{S-S}(f) = \left(\frac{\Delta \varphi_{nl}}{2\pi^2 f T_{rt} \nu_0} \right)^2 S_I(f)$	Phase shift (timing jitter) caused by nonlinear refractive index (intensity dependant)
Slow-saturable absorber (SSA)	$S_{\Delta t}^{SSA}(f) = \left(\frac{sd\Delta t/ds}{2\pi f T_{rt}} \right)^2 S_I(f)$	Temporal shift induced by the fact that the leading part of the pulse is more attenuated by the saturable absorber than the trailing part
Total	$S_{\Delta t}^{tot} = \sum_i S_{\Delta t}^i$	i stands for the supra-indexes corresponding to all the physical effects above

Equations from R. Paschotta, Appl. Phys. B, 79 163-173, 2004

Why DPSSL is low noise?



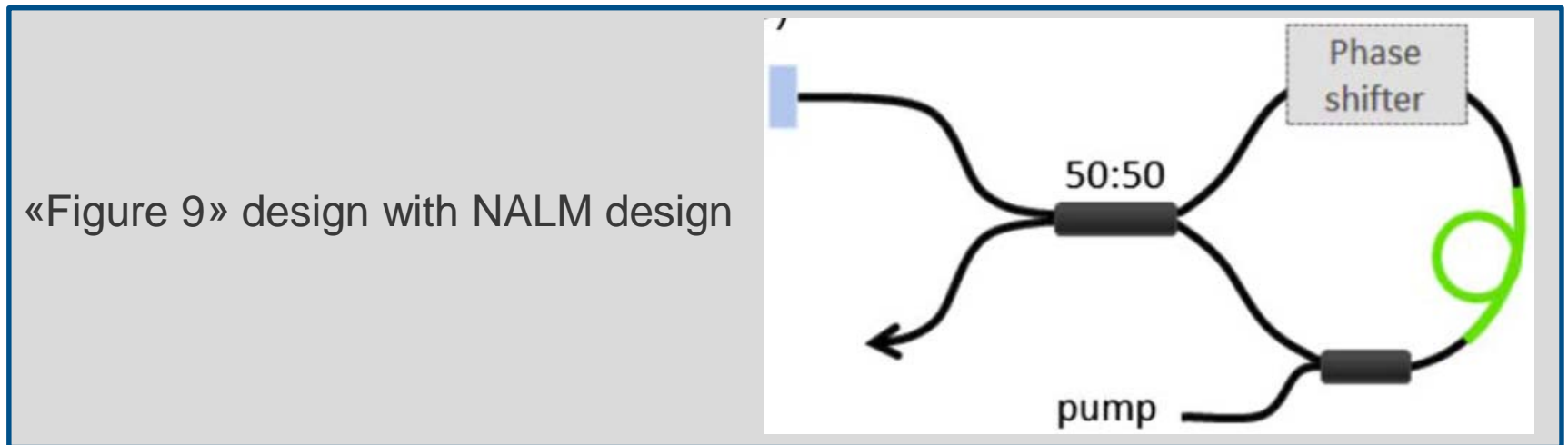
- **Low gain**
- **Typically producing solitons**
- **Large pulse energy**
- **Typical repetition rates: 40 MHz – 1 GHz**



Low quantum noise

And fiber lasers?

- :: Large number of different architectures
- :: Best performances obtained with nonlinear amplifying loop mirror (NALM) architecture thanks to instantaneous saturable absorber (Kerr lens like)

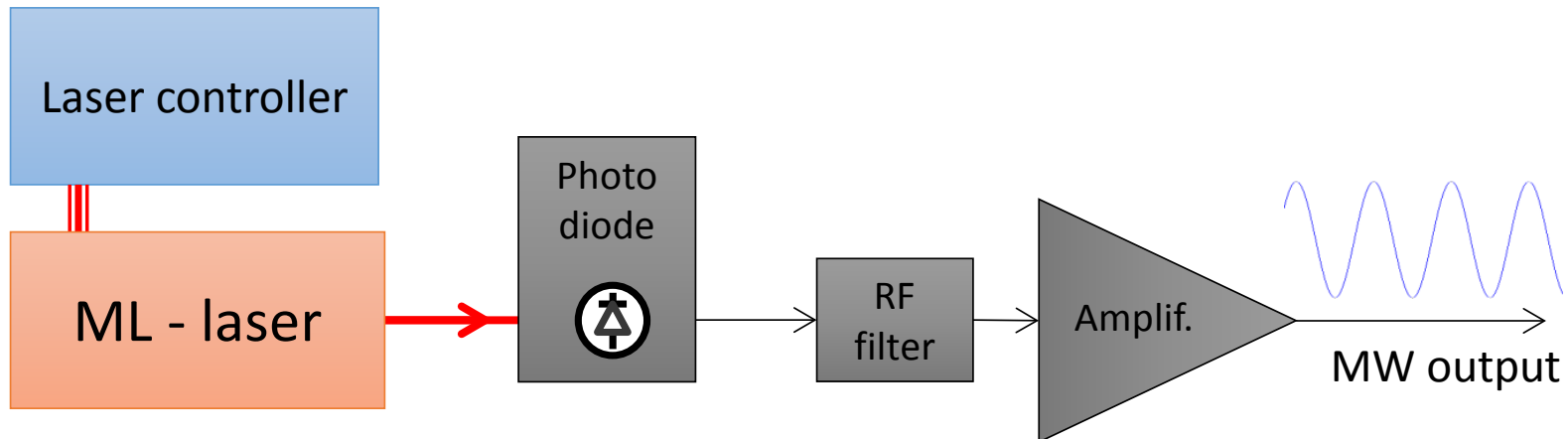


- :: Compact and robust design
- :: Limited pulse energy and relatively high intracavity losses
- :: Timing jitter larger than solid-state lasers

Figure from J. Kim and Y. Song, Adv. In Opt. And Phot. 8 465-540, 2016

Optical pulse timing jitter and microwave phase noise

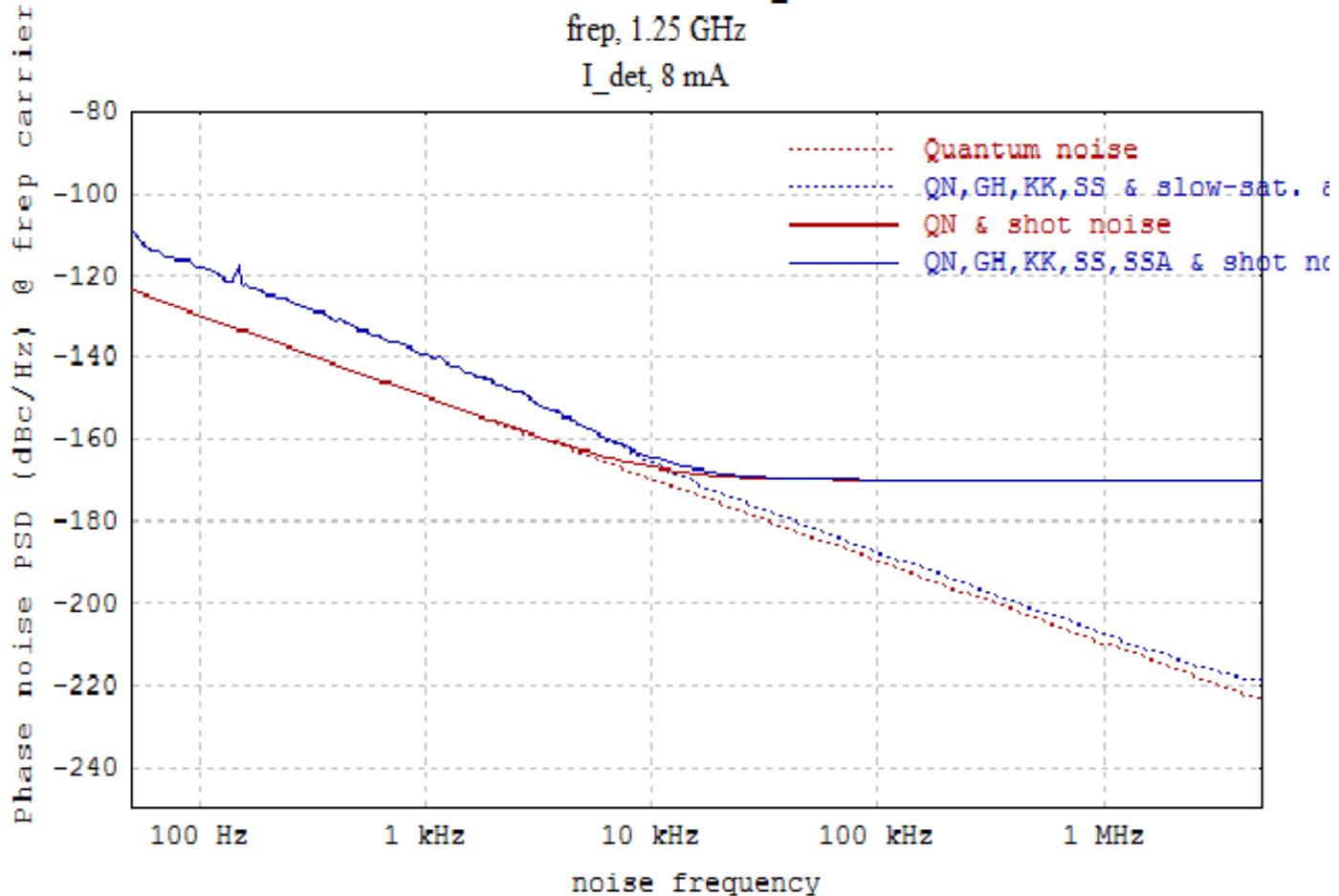
Single sideband phase noise	$L(f) = 10 \log(4\pi^2 f_c^2 S_{\Delta t}(f)), \text{ dBc/Hz}$
Shot noise associated single sideband phase noise	$L_{SN}(f) = 10 \log\left(\frac{q}{2 I_{av}}\right), \text{ dBc/Hz}$



Timing jitter modelling

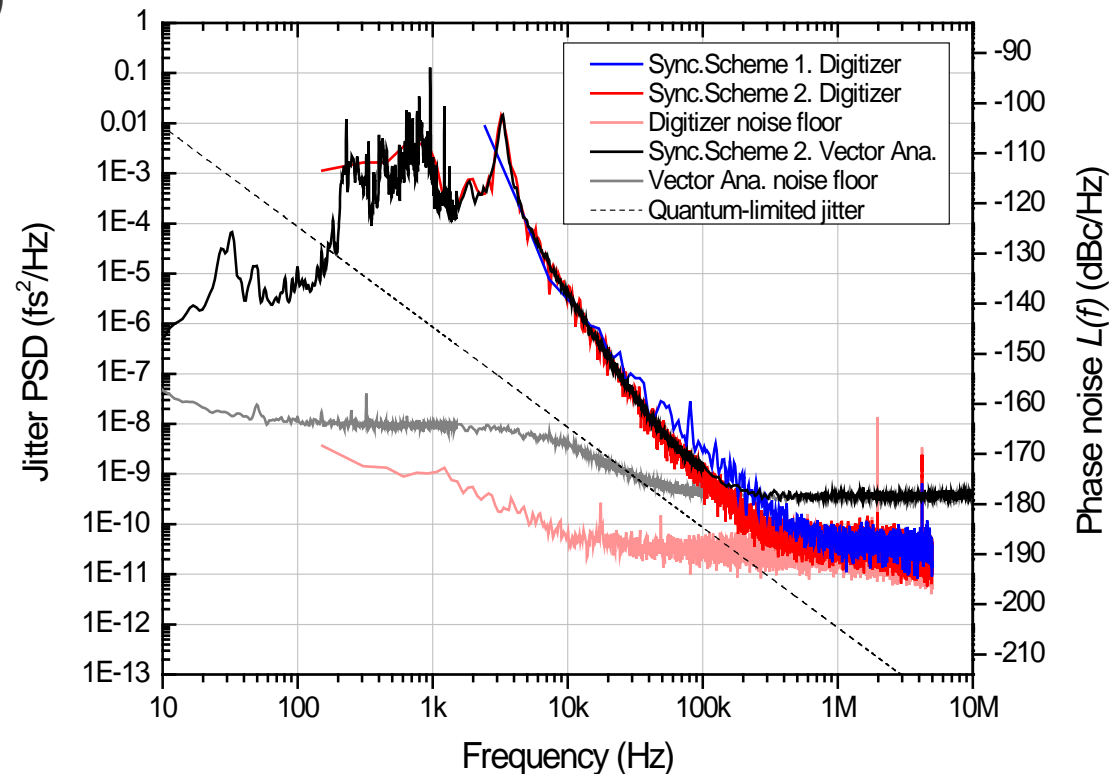
RF Phase Noise Spectrum

frep, 1.25 GHz
I_det, 8 mA



Timing jitter of DPSSL

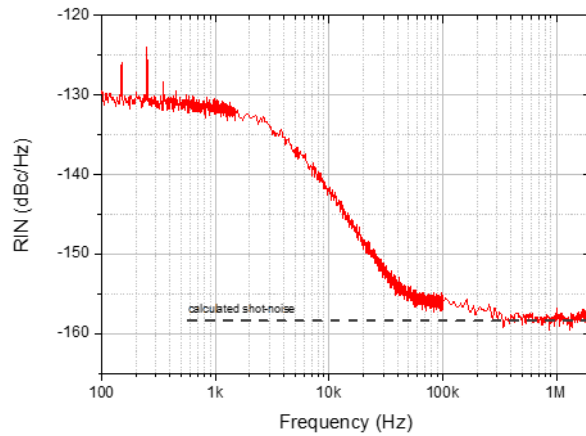
- :: Optical pulse train jitter of 100-MHz rep rate 1556 nm DPSSL
- :: Optical measurement (no optical-to-electrical conversion), 83 as (integrated from 10 kHz to 50 MHz)



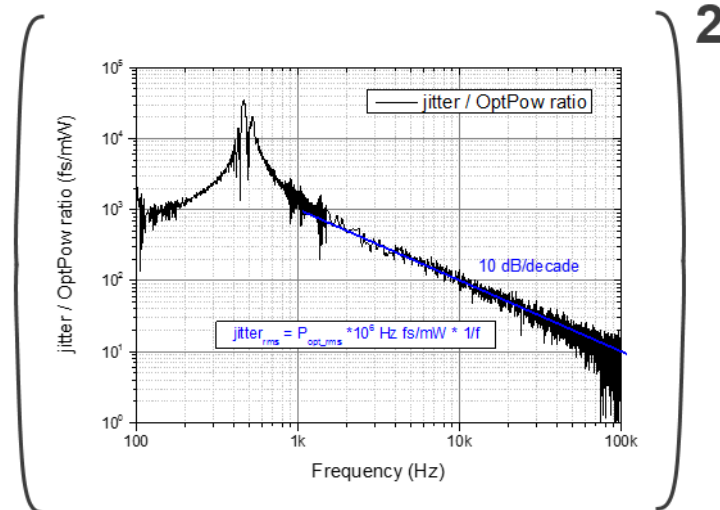
E. Portuondo-Campa et al., Opt. Lett. 38, 2650-2653 (2013)

Investigation of extra-noise source

Measured Amplitude Noise

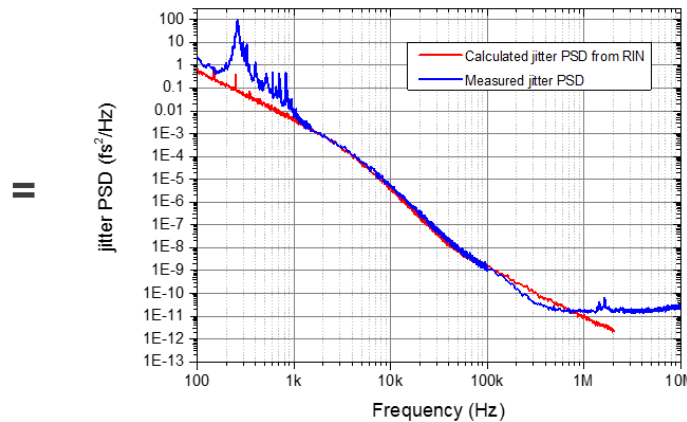


Measured AM to PM conversion



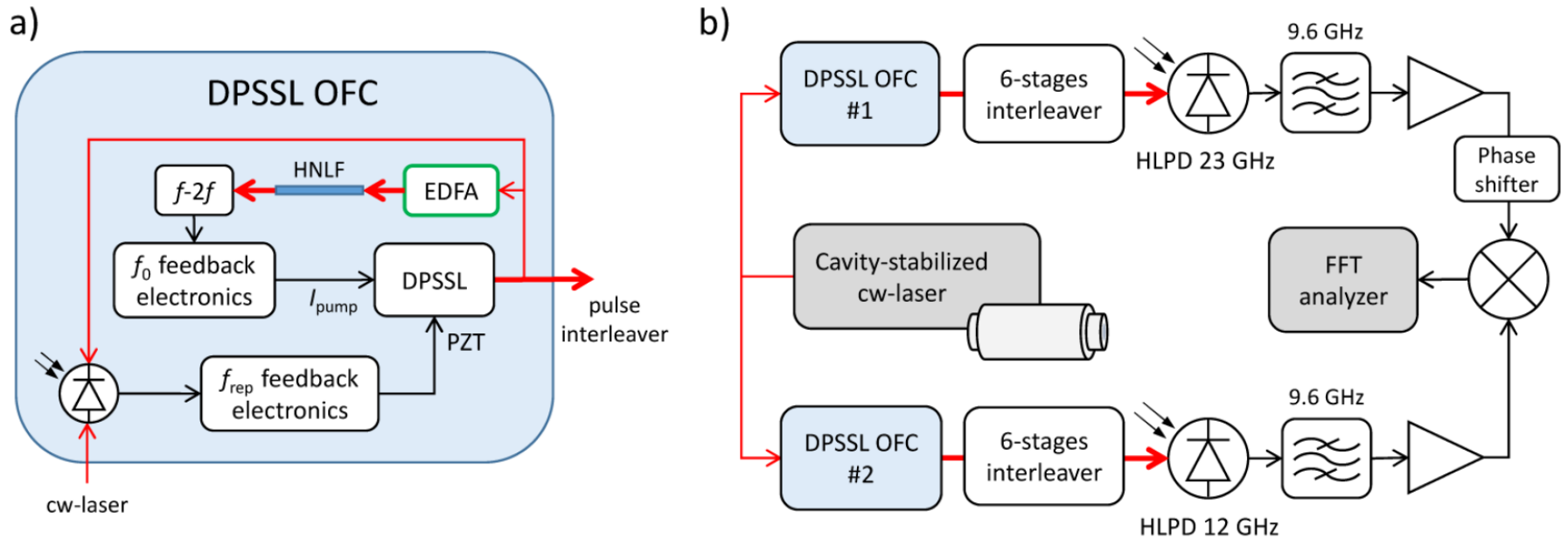
X

Calculated Jitter power spectral density



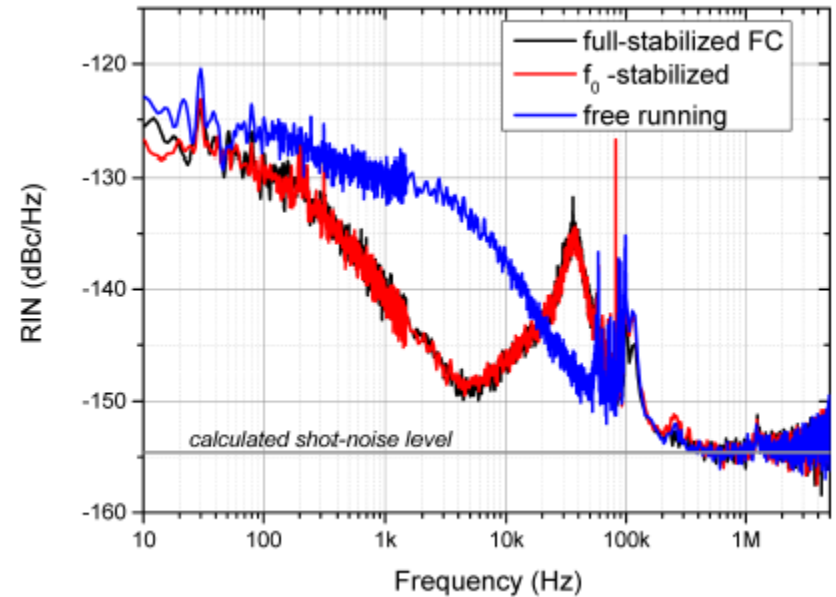
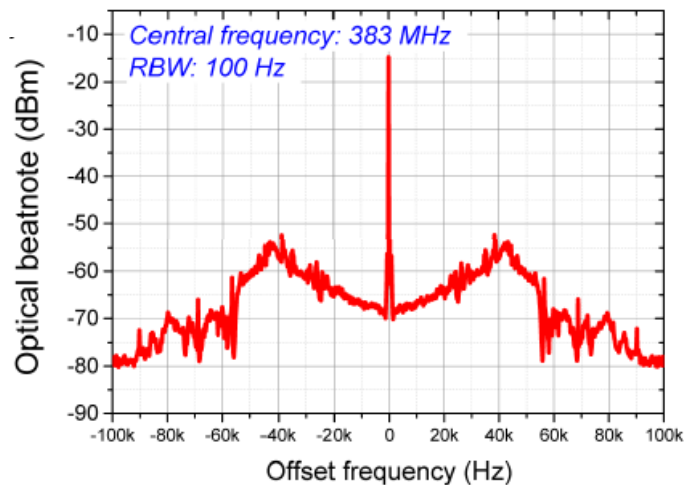
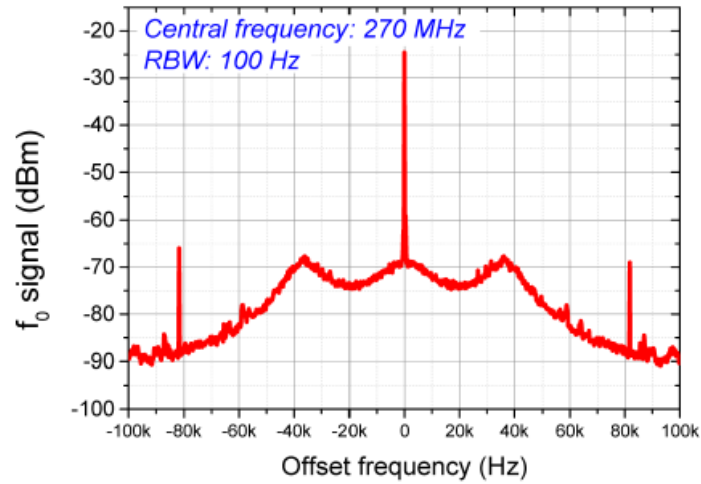
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Ultrapure microwave generation



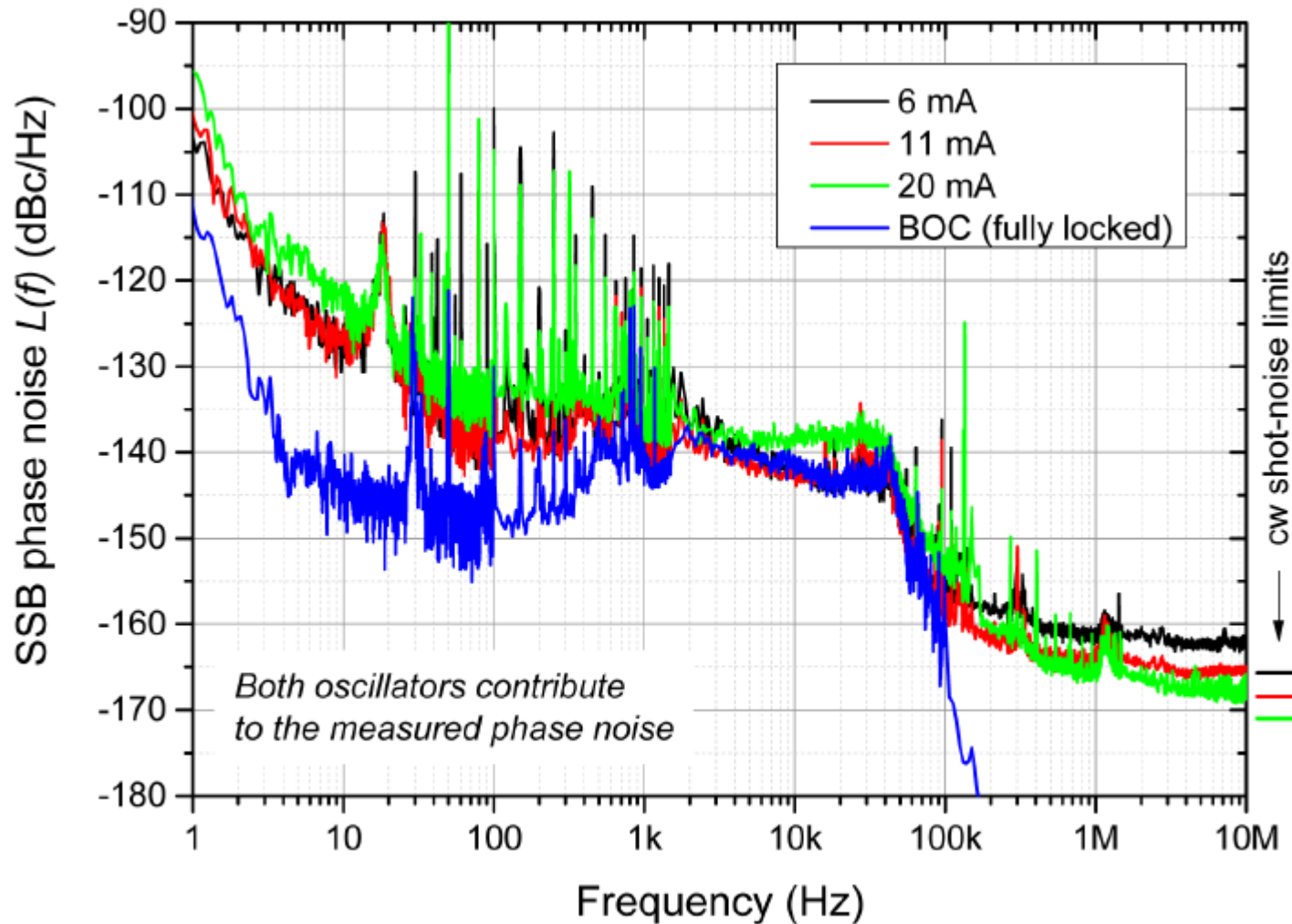
E. Portuondo-Campa *et al.*, Opt. Express 23, 32441-32451 (2015)

Stabilized 100MHz DPSSL comb



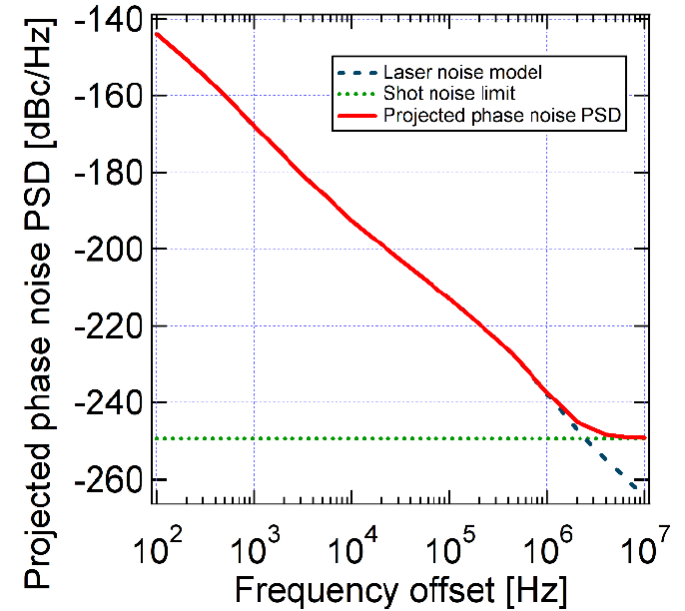
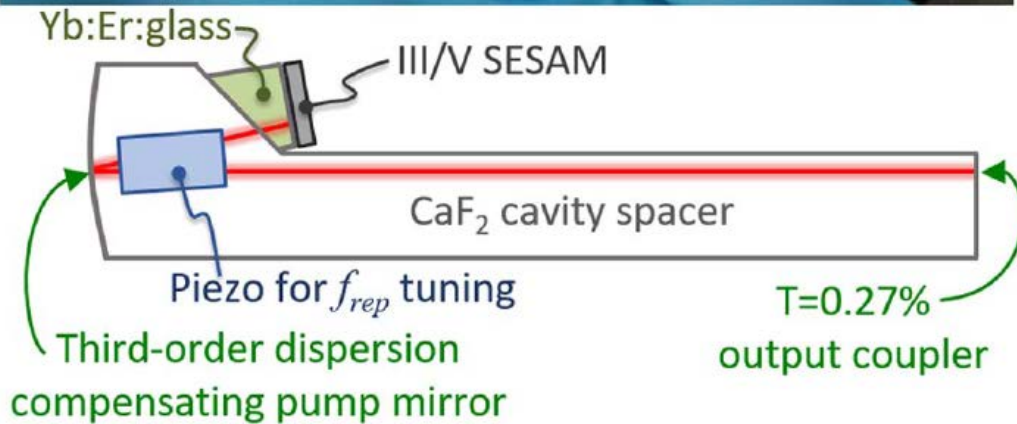
E. Portuondo-Campa *et al.*, Opt. Express 23, 32441-32451 (2015)

Microwave phase noise and pulse timing jitter



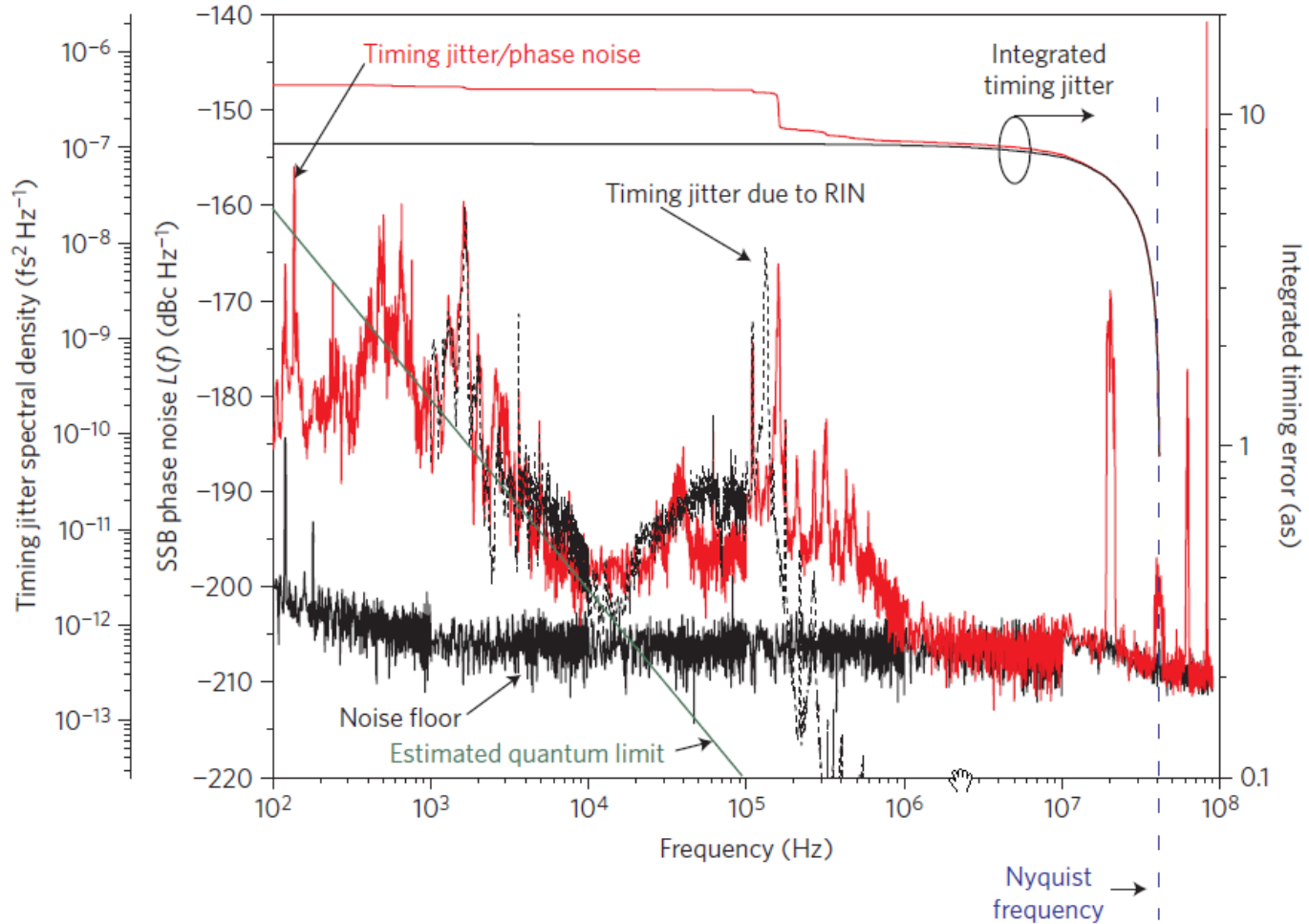
E. Portuondo-Campa *et al.*, Opt. Express 23, 32441-32451 (2015)

Ultralow noise 1GHz MLL based on monolithic design



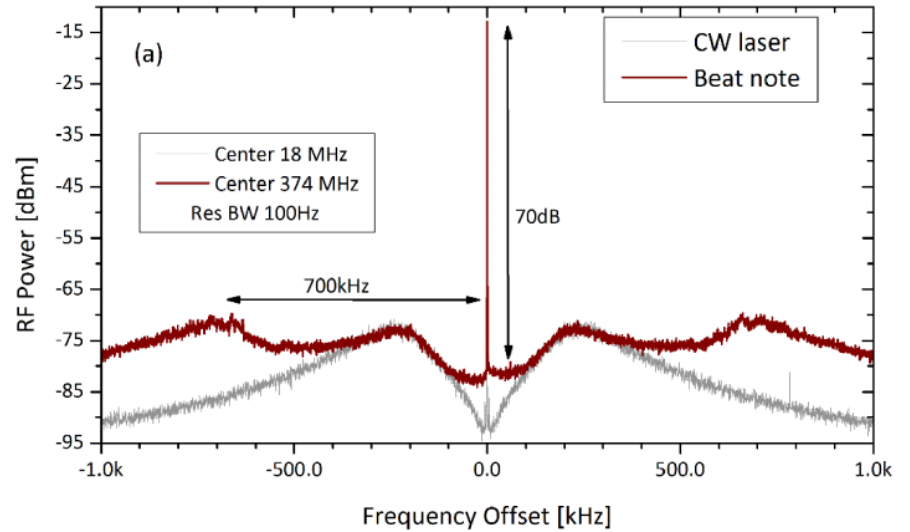
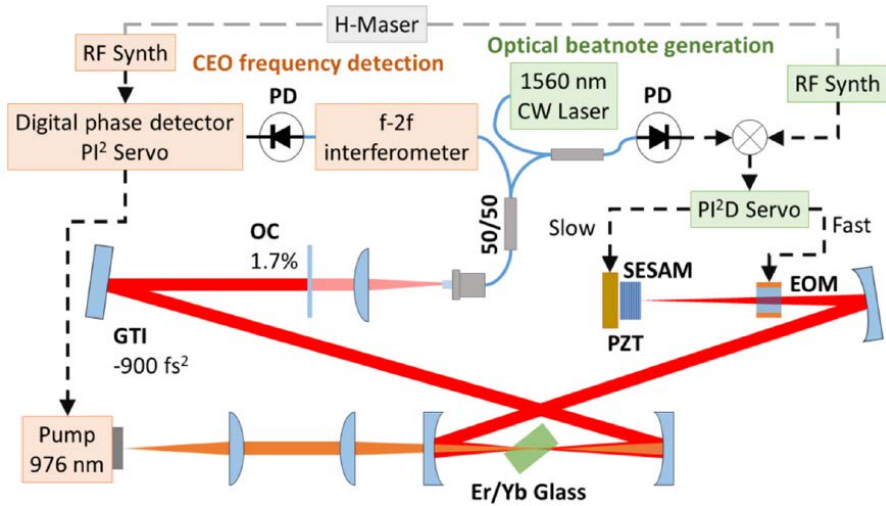
T. D. Shoji *et al.*, *Optica* 3, 995-998 (2016)

100MHz Kerr lens Titanium Sapphire pulse timing jitter



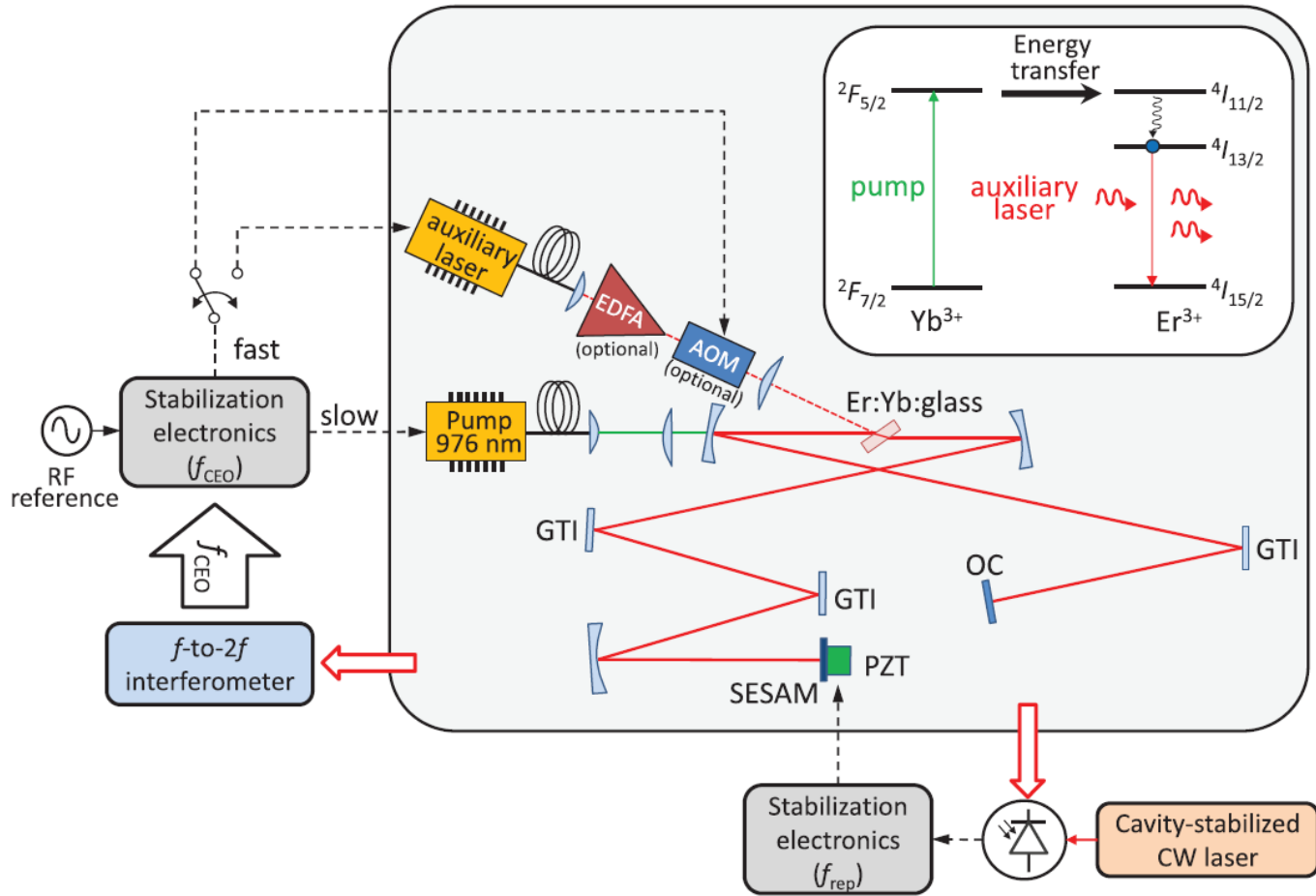
A. J. Benedick *et al.*, Nat. Phot. 6, 97-100 (2012)

EOM in DPSSL



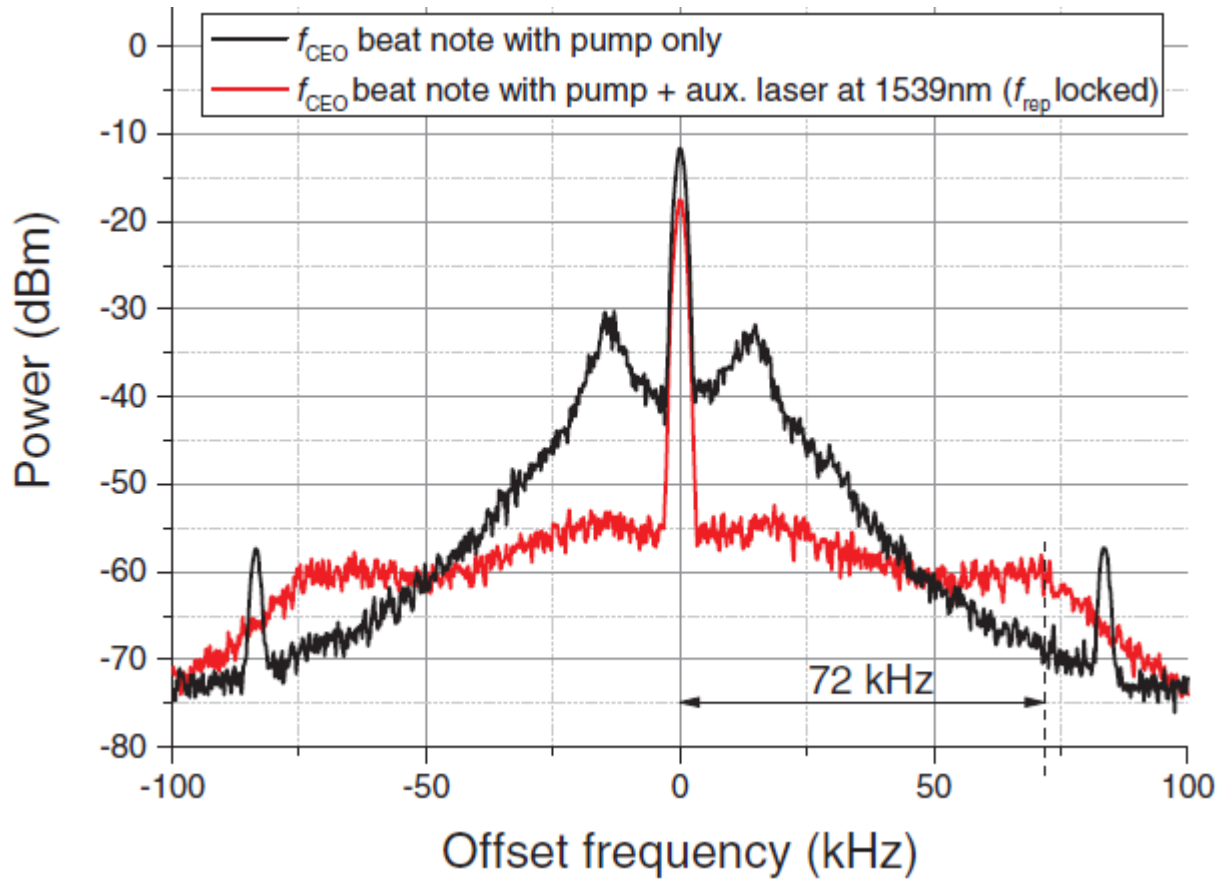
N. Torcheboeuf *et al.*, Opt. Express 25, 2215-2220 (2017)

Power / CEO stabilization via gain control



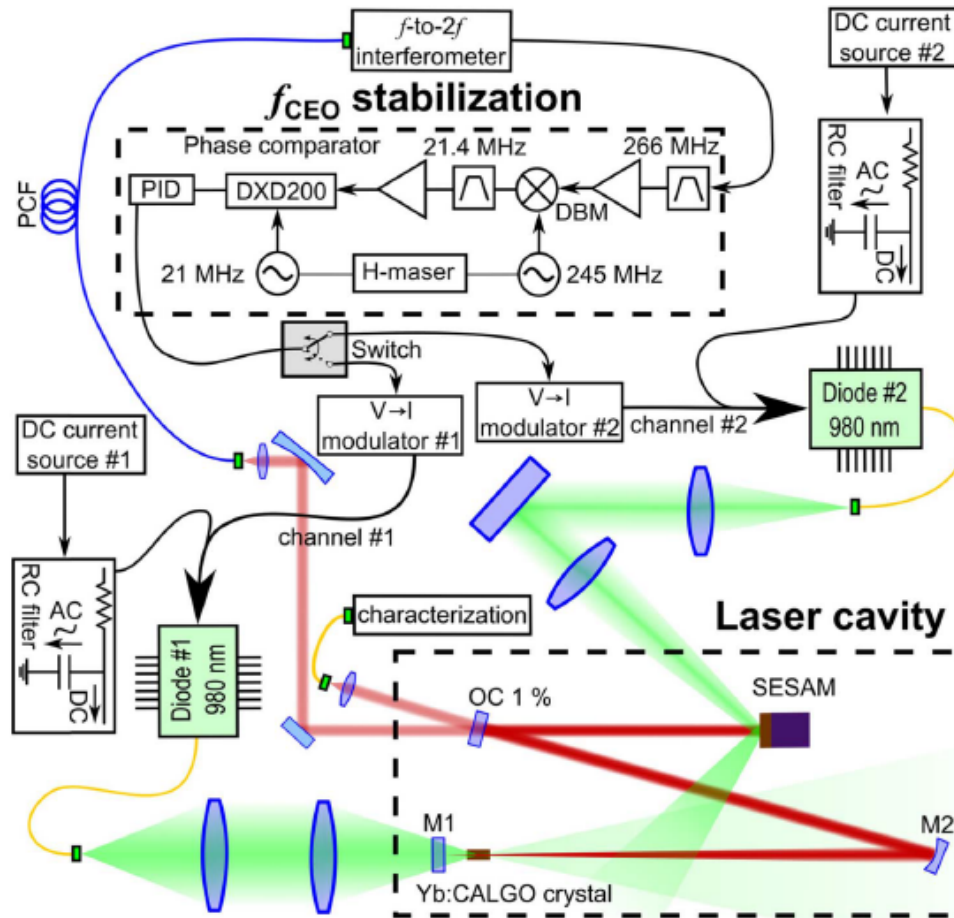
L. Karlen *et al.*, Opt. Letters 41, 376-379 (2016)

Power / CEO stabilization via gain control



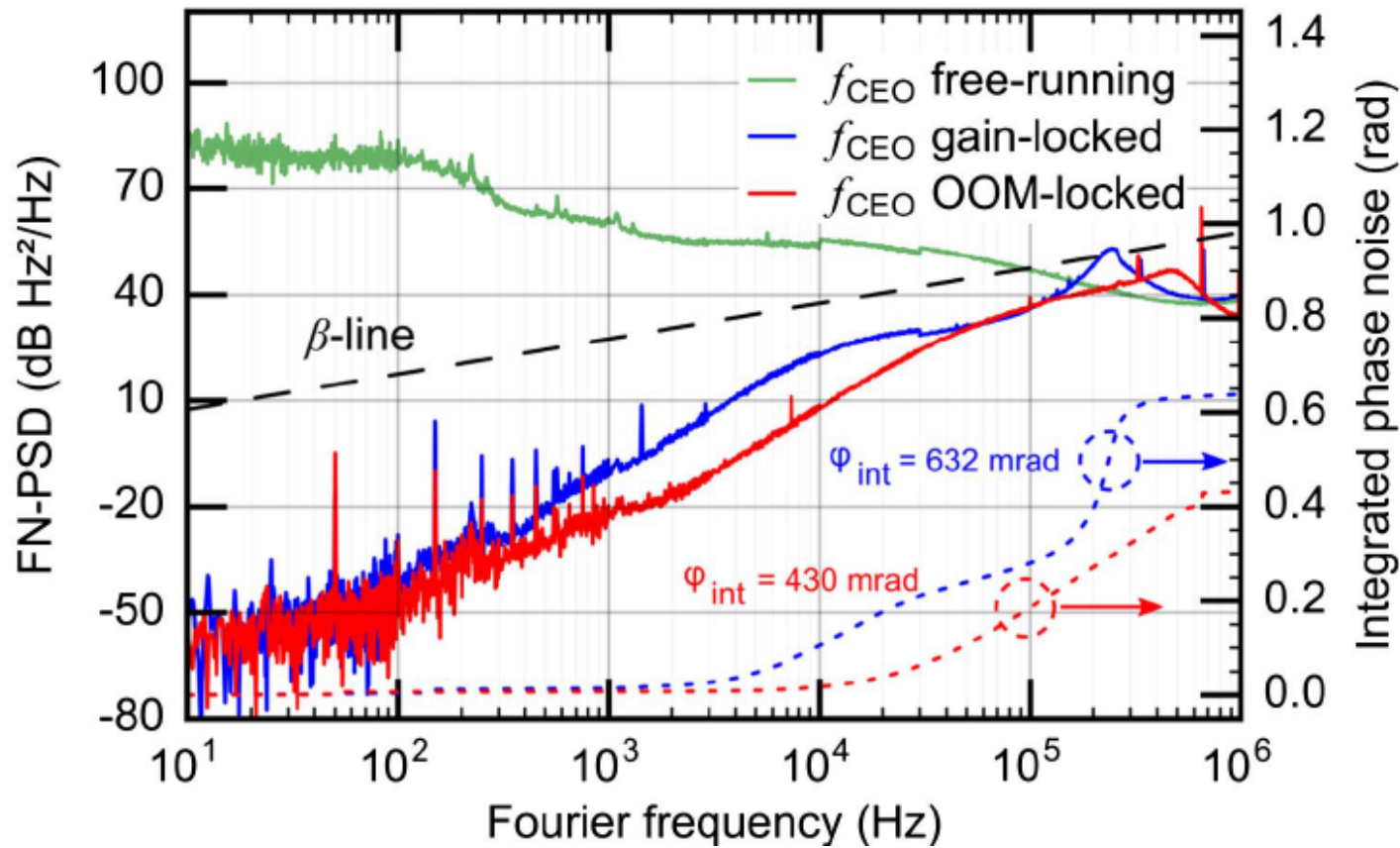
L. Karlen *et al.*, Opt. Letters 41, 376-379 (2016)

Optical actuation on SESAM



S. Hakobyan *et al.*, Opt. Letters 42, 4651-4654 (2017)

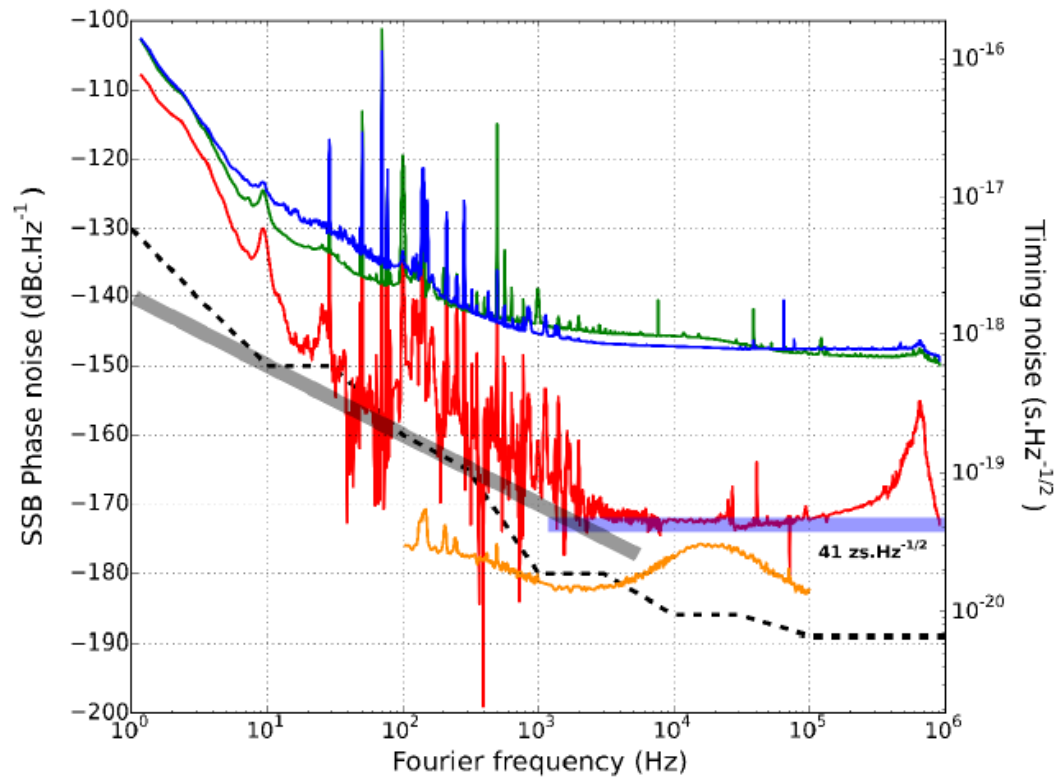
Optical actuation on SESAM



S. Hakobyan *et al.*, Opt. Letters 42, 4651-4654 (2017)

Fiber laser with fast actuators

- :: EOMs for frep and CEO control allow MHz range locking bandwidths
- :: Efficient implementation in loss tolerant fiber lasers for tight locking



X. Xie *et al.*, Nat. Phot. 11, 44-47 (2017)

Conclusions and outlook

- :: fs modelocked lasers are exquisitely good oscillators
- :: Fundamental and technical noise sources in such lasers are well known and understood
- :: New laser designs are under development
- :: Still room to push for improved laser performances via reducing noise sources and optimized stabilization
- :: Better photonics-based microwave oscillators will be realized

Reference articles

- :: J. Kim and Y. Song, Adv. In Opt. And Phot. 8 465-540, 2016
- :: N. R. Newbury and W. C. Swann, JOSA B 24 1756-1771, 2007
- :: R. Paschotta, Appl. Phys. B, 79 163-173, 2004

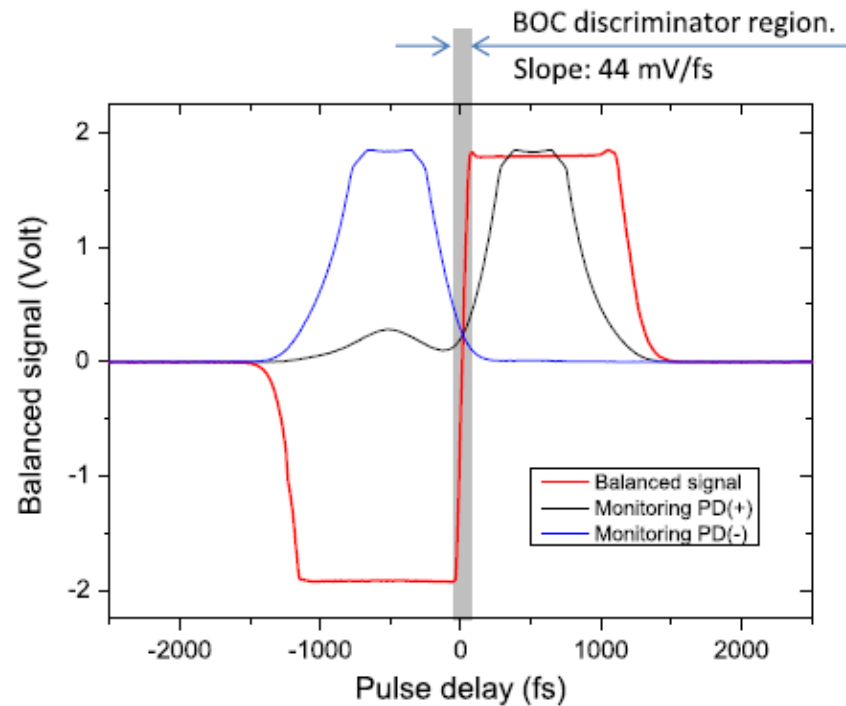
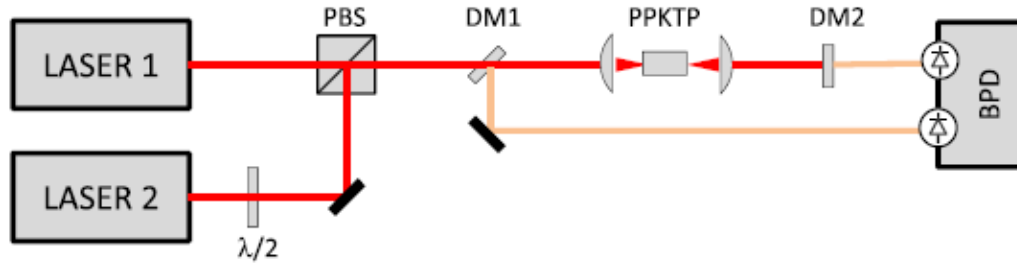
Thank you for your attention!

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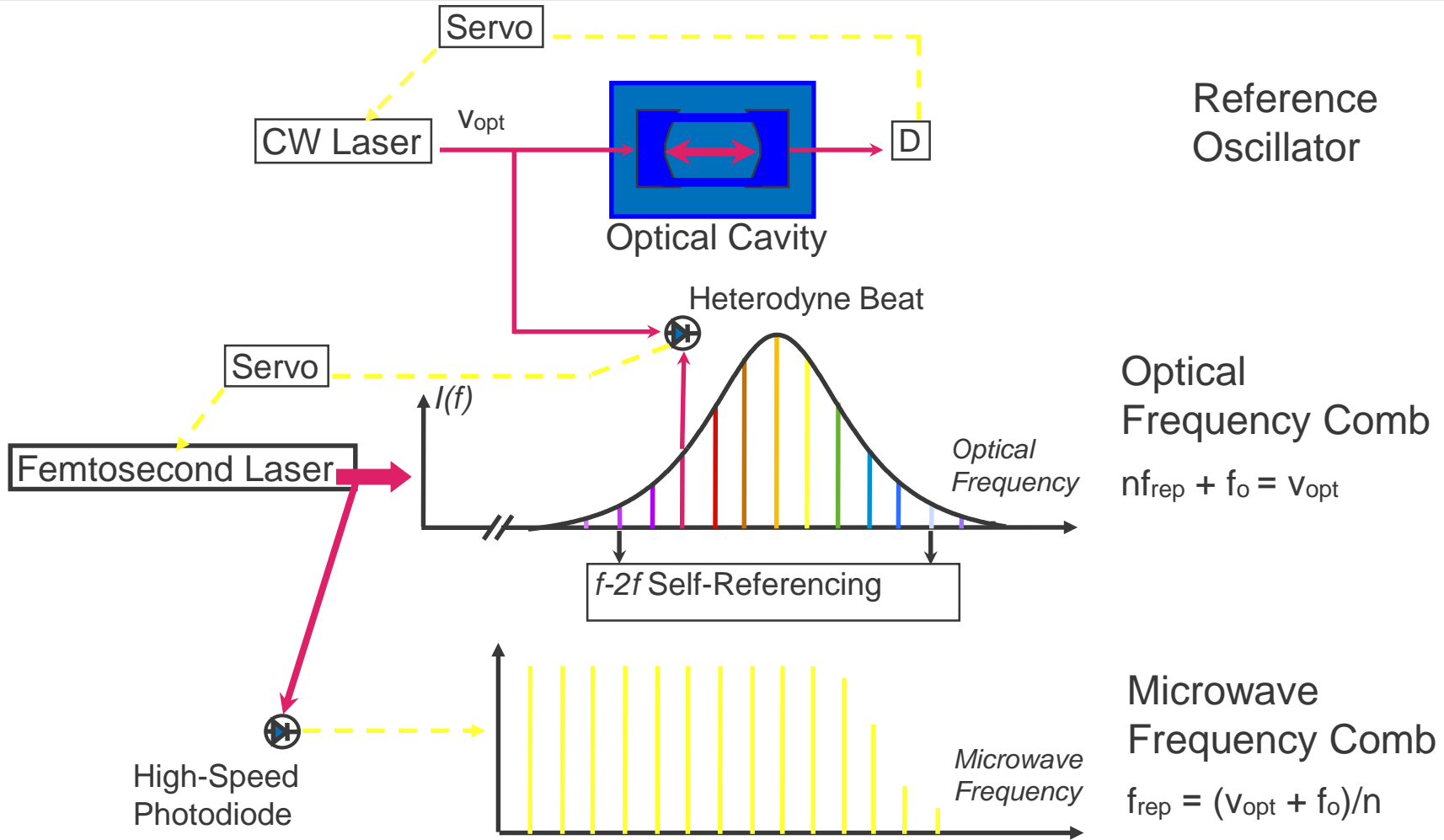
www.csem.ch

BOC technique



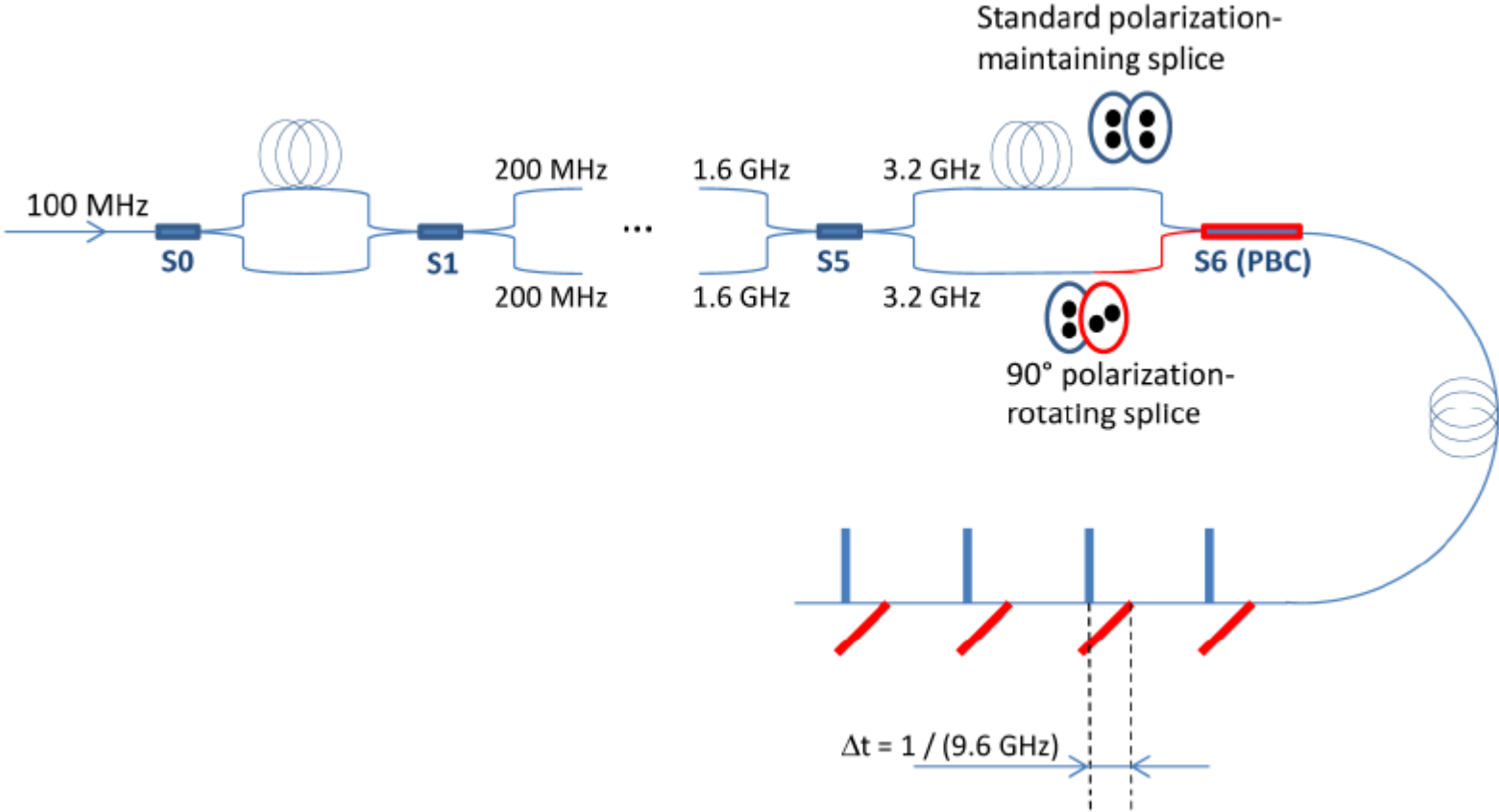
E. Portuondo-Campa et al., Opt. Lett. 38, 2650-2653 (2013)

Ultrapure microwave generation



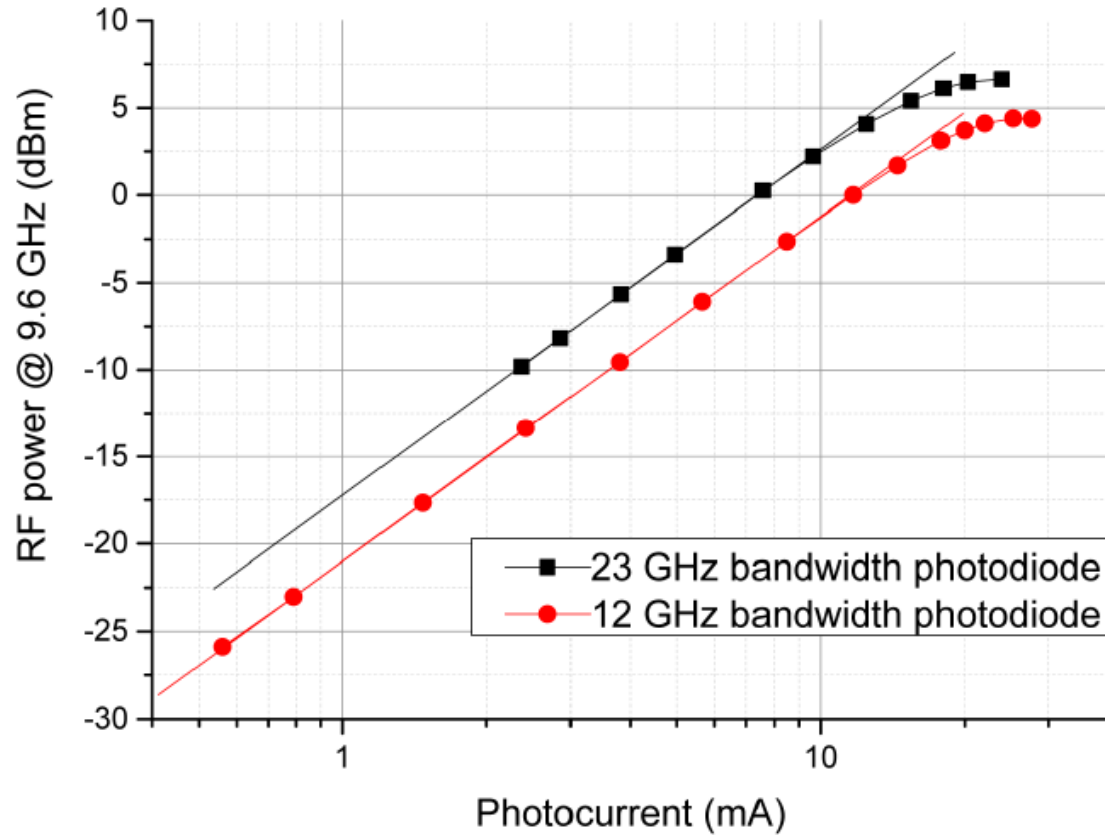
Slide from S. Diddams, NIST

Augmentation du taux de répétition du train d'impulsion



E. Portuondo-Campa *et al.*, Opt. Express 23, 32441-32451 (2015)

Photocourant et sa saturation



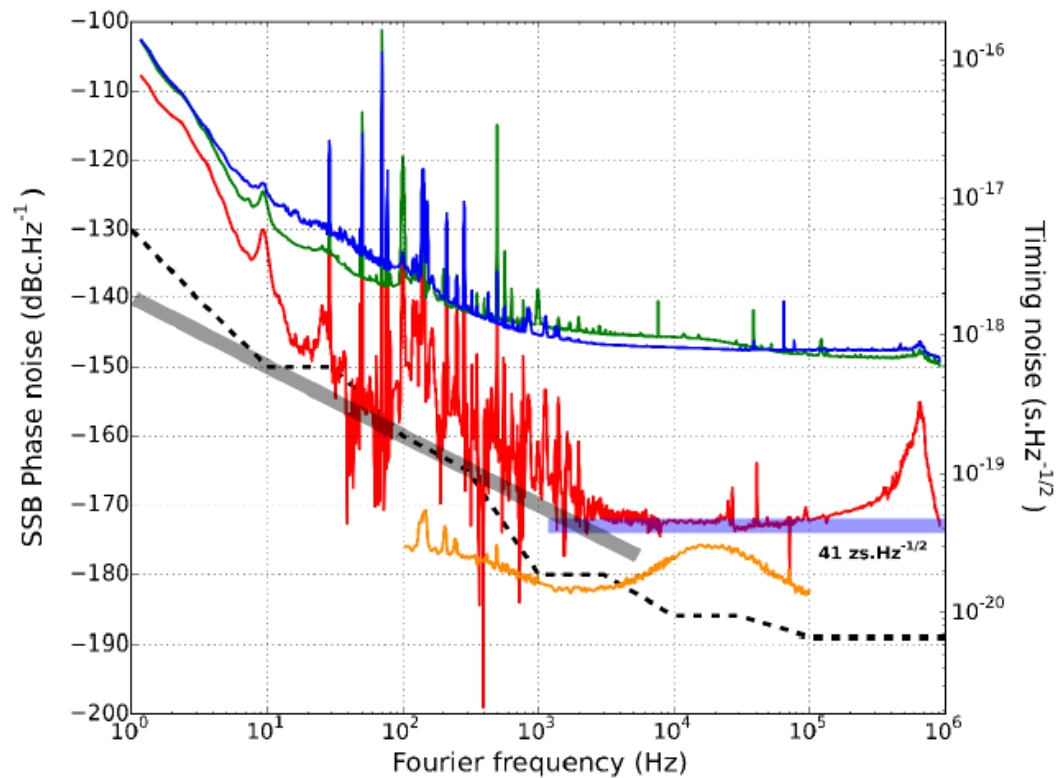
E. Portuondo-Campa *et al.*, Opt. Express 23, 32441-32451 (2015)

Cavité de référence



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X. Xie *et al.*, Nat. Phot. 11, 44-47 (2017)